# Anisotropic Flow; ad nauseam

**Raimond Snellings** 



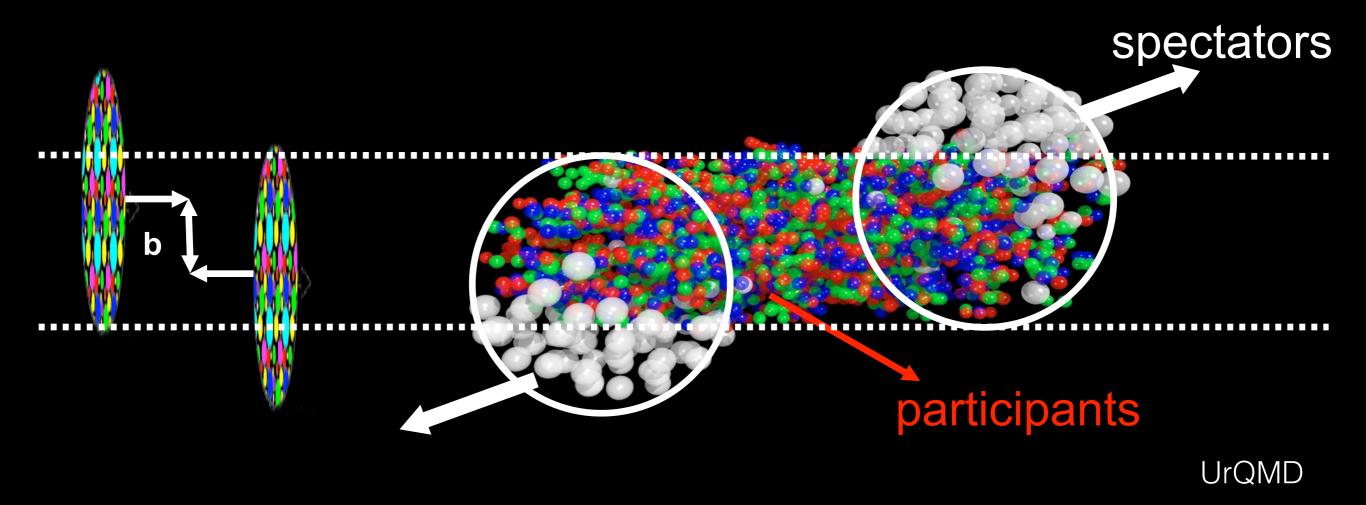
NI<mark>KH</mark>EF

4<sup>th</sup> International Symposium on Non-equilibrium Dynamics 30-08 — 05-09-2015 Giardini Naxos, Sicily,Italy

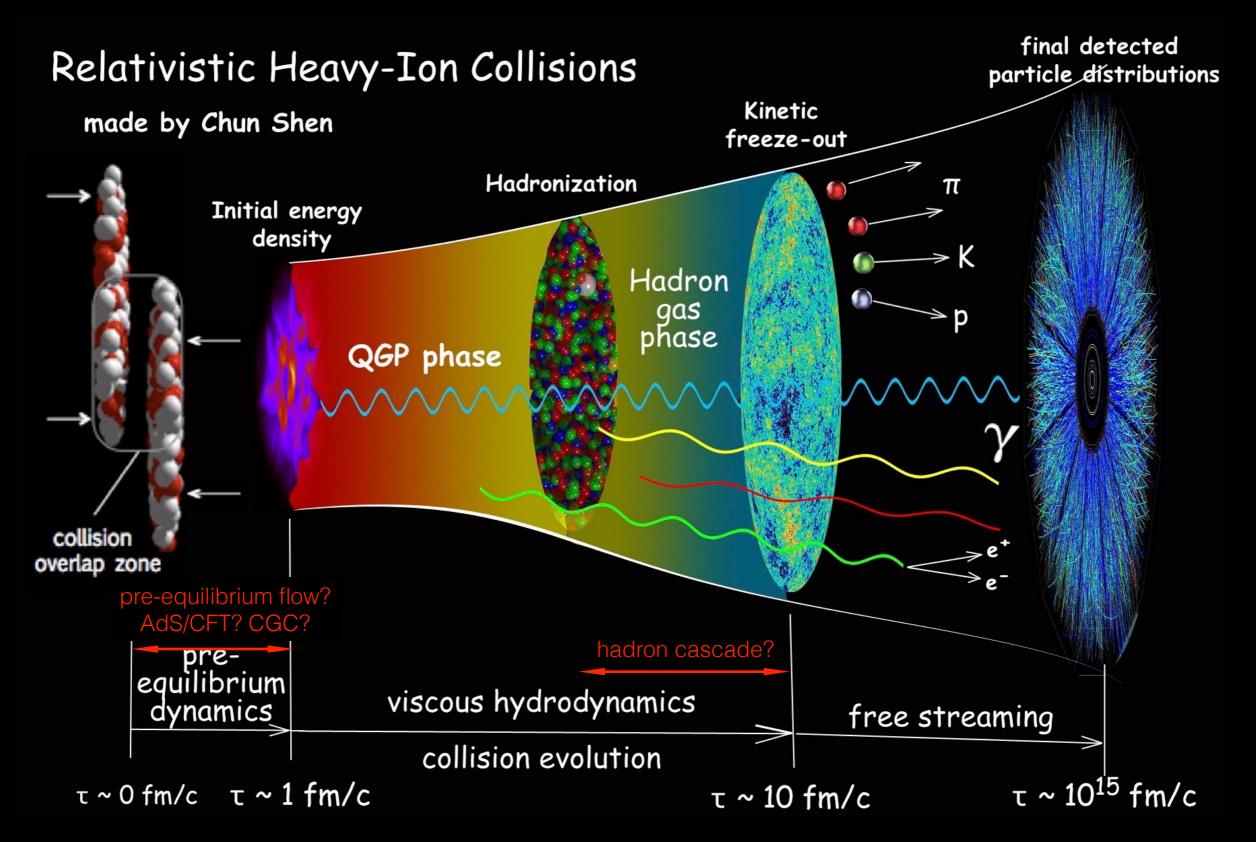
## Outline

- initial spatial distributions and the response of the system
- integrated elliptic flow
  - EoS, Knudsen number and  $\eta/s$
- p<sub>t</sub>-differential elliptic flow
  - identified particles
- other harmonics
- what do we measure?
- v<sub>n</sub> probability distributions
- rapidity dependence

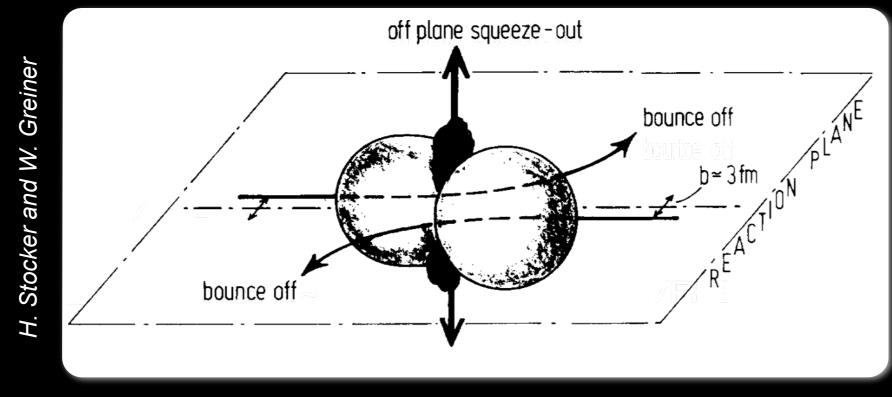
## A Heavy-Ion Collision



### Our current picture



## A long long time ago



#### Jean-Yves Ollitrault; PRD 46 (1992)

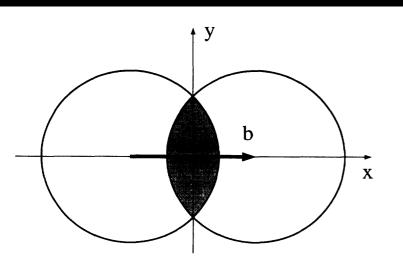
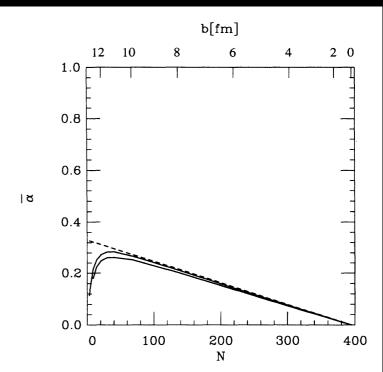
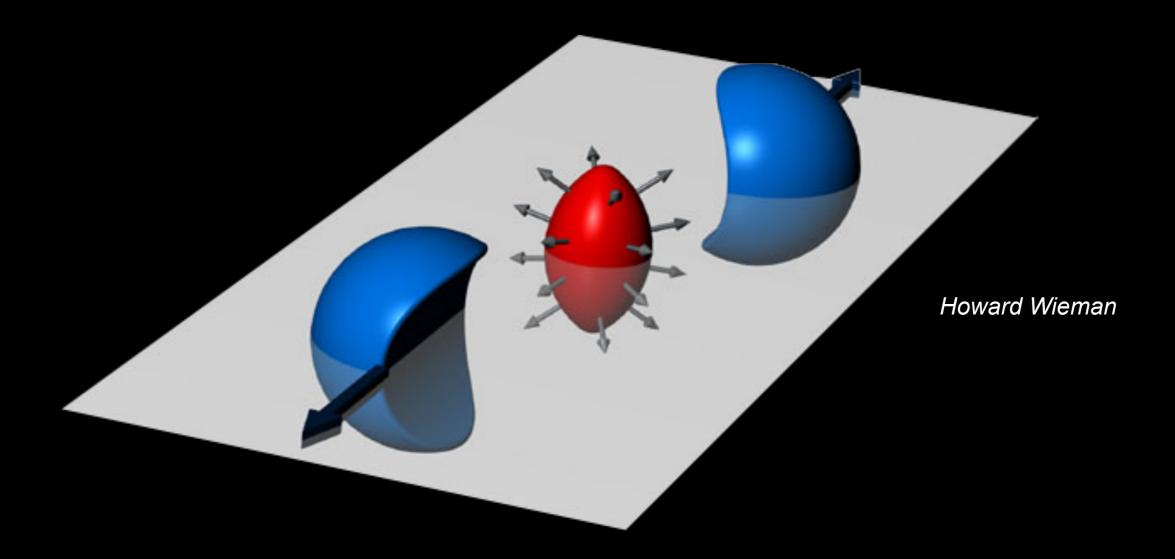


FIG. 1. Peripheral collision viewed in the transverse plane. **b** is the impact parameter. The shaded area corresponds to the region where particles are created in the central rapidity region. Outside this region is the vacuum.

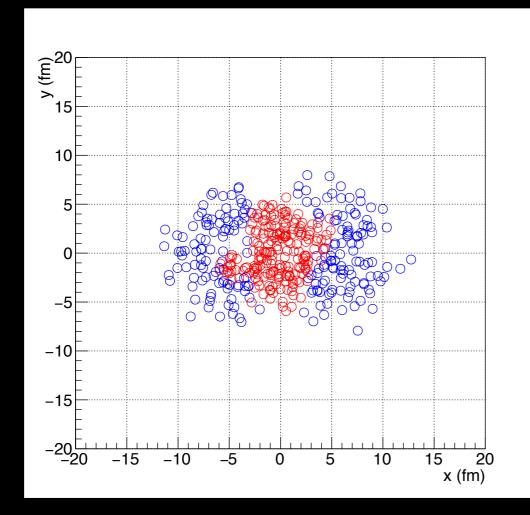


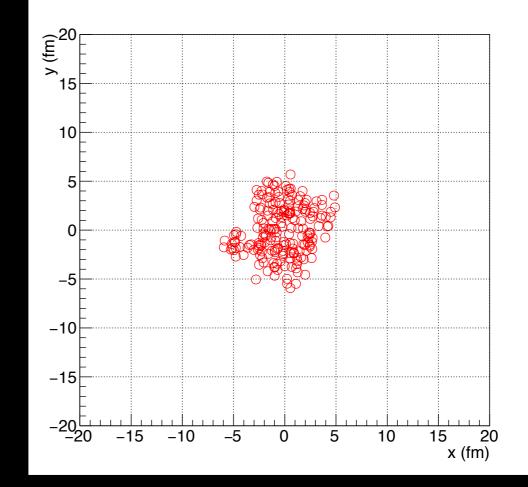
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### The Classical Picture

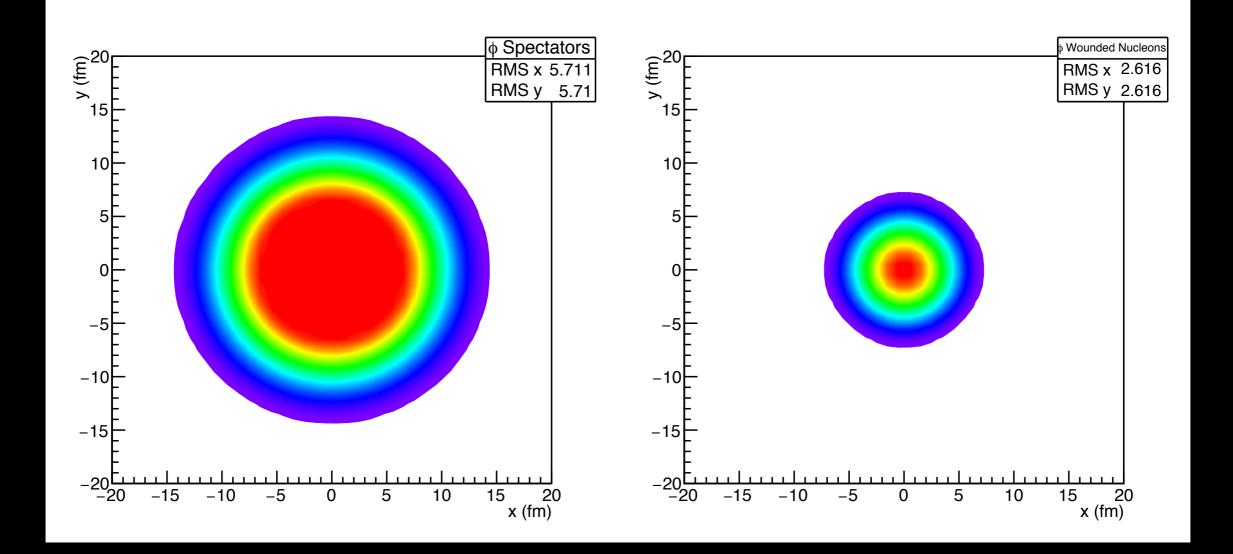


## A Single Collision

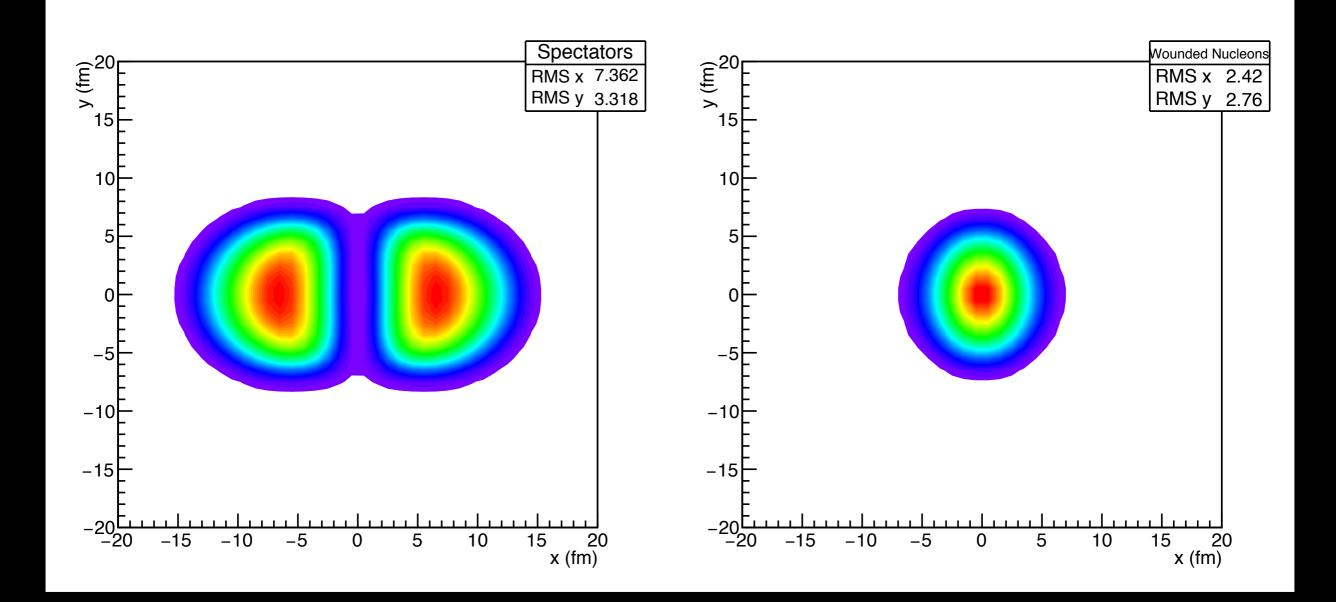




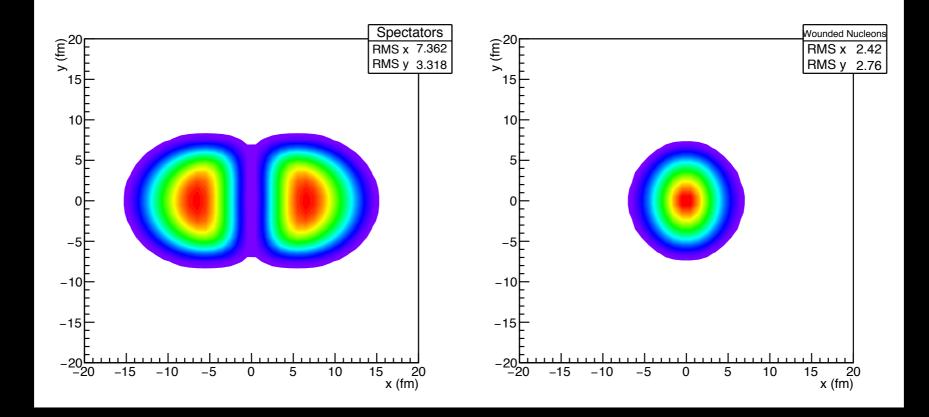
### Many Collisions in the Lab Frame

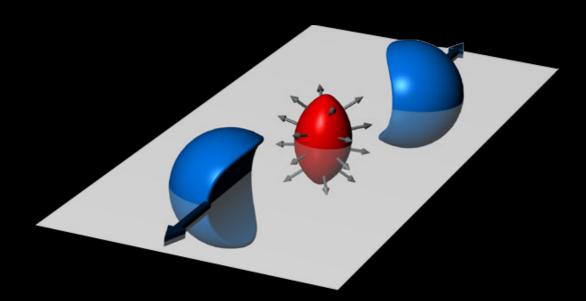


### Many Collisions versus the Reaction Plane



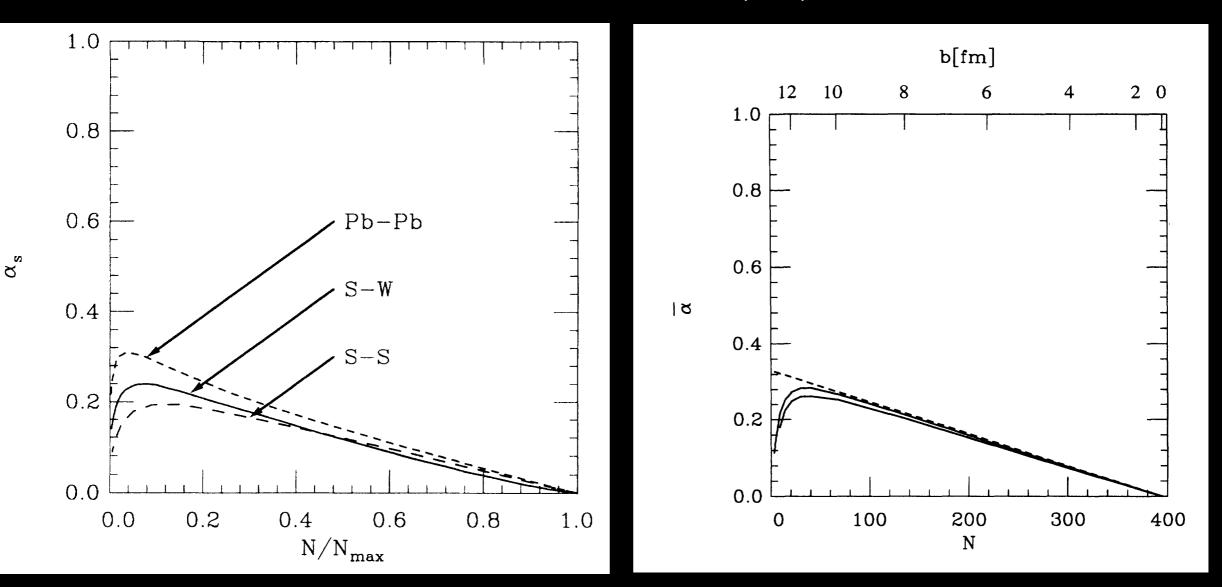
## A long long time ago





## A long long time ago

Jean-Yves Ollitrault; PRD 46 (1992)

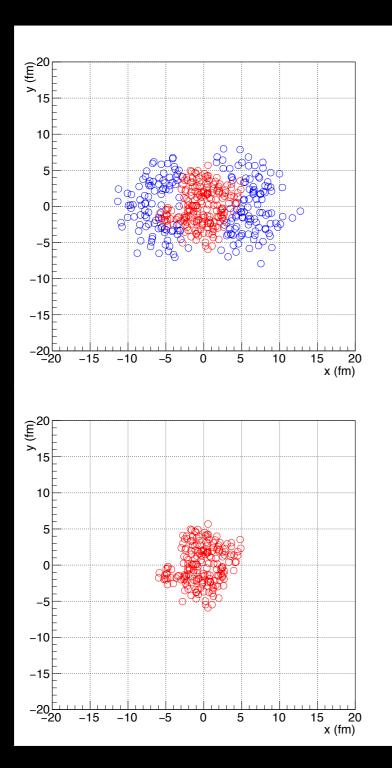


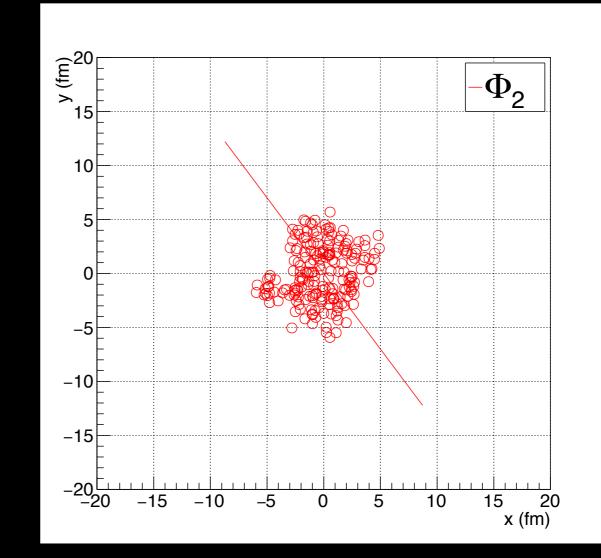
 $v_n \propto \varepsilon_n$ 

sensitive to the EoS

11

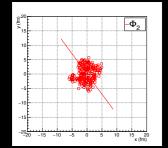
## Symmetry Plane



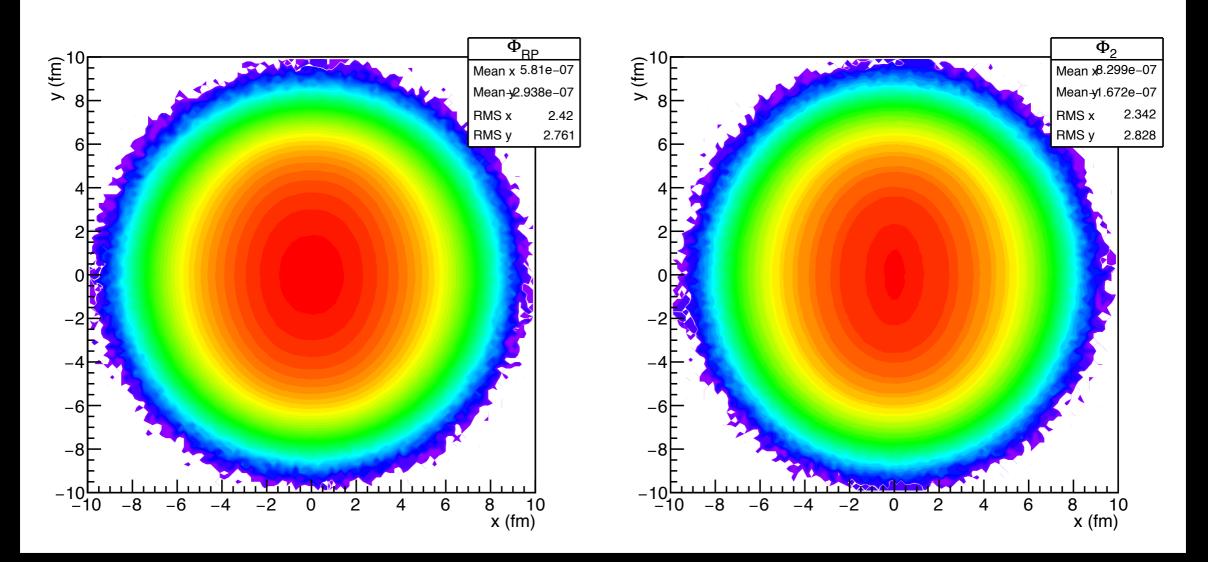


Using the particles produced we (experimentalists) determine, due to the fluctuations, a symmetry plane which is different than the Reaction Plane

 $v_n \propto \varepsilon_n$ 



## Symmetry Planes

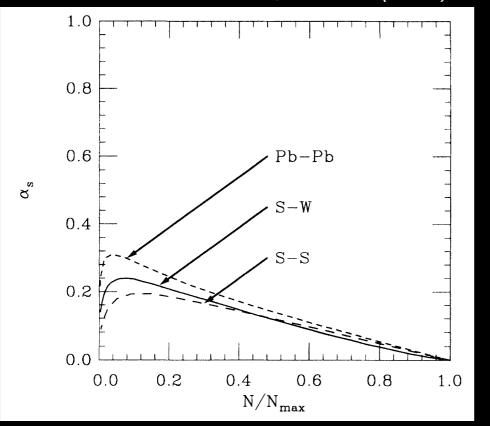


The asymmetry of the system is larger versus this symmetry plane

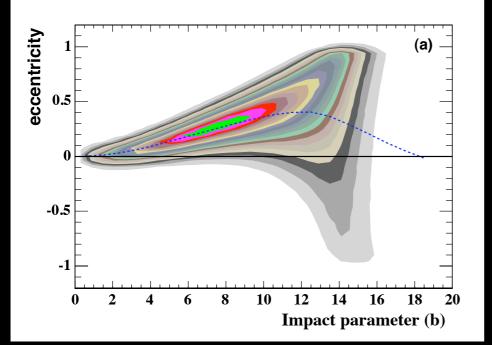
 $v_n \propto \varepsilon_n$ 

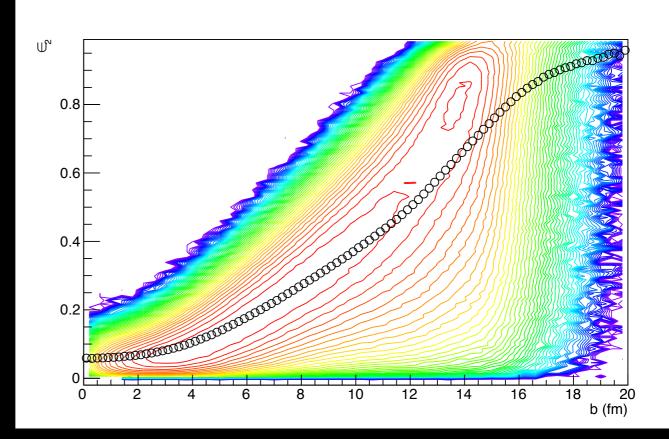
### Fluctuations

Jean-Yves Ollitrault; PRD 46 (1992)



*Mike Miller, RS nucl-ex/0312008 (2003)* 



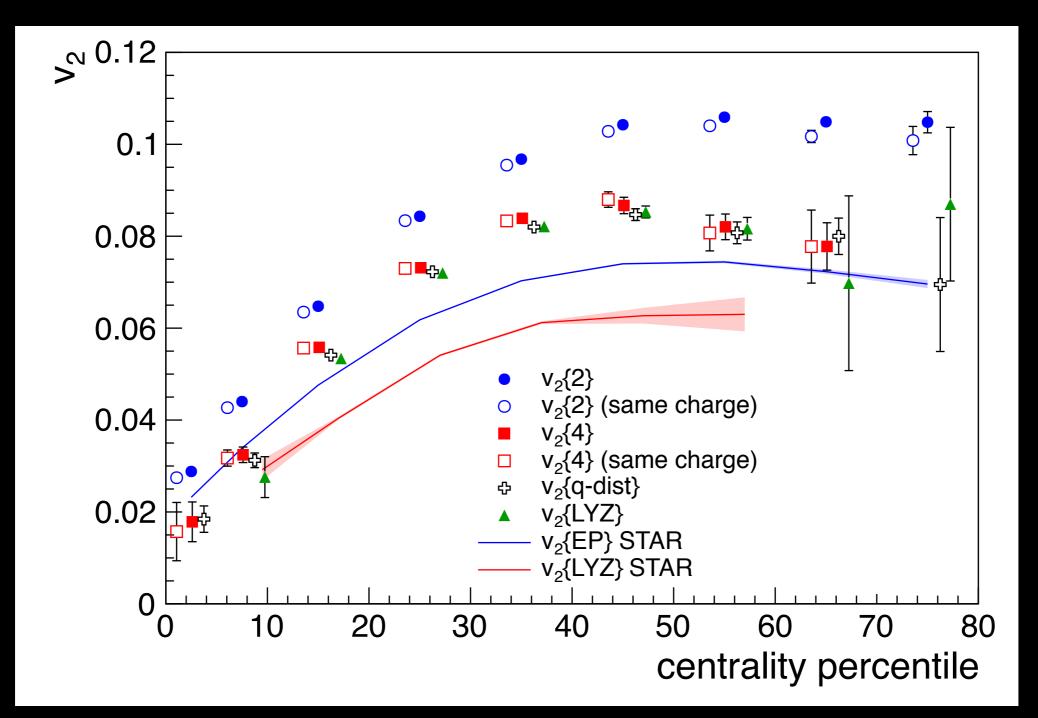


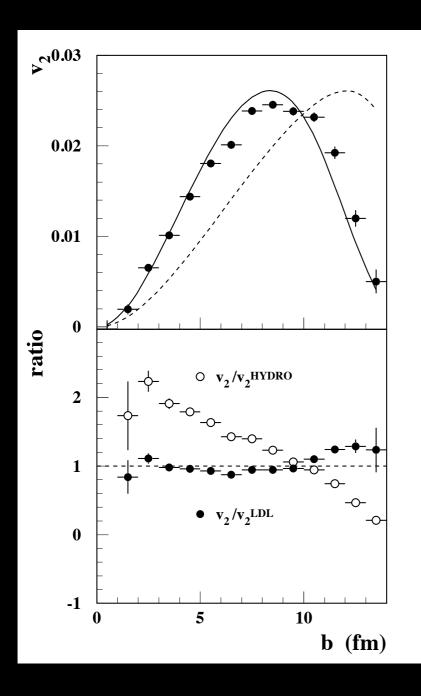
The asymmetry is larger and even non-zero for perfectly central collisions This asymmetry in coordinate space is though to be responsible, due to e.g. final state interactions, for the observed anisotropy in particle production

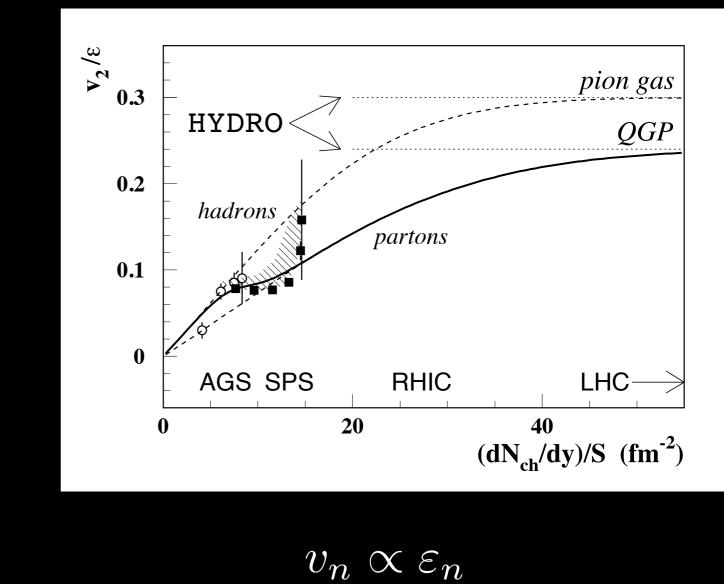
$$v_n \propto \varepsilon_n$$

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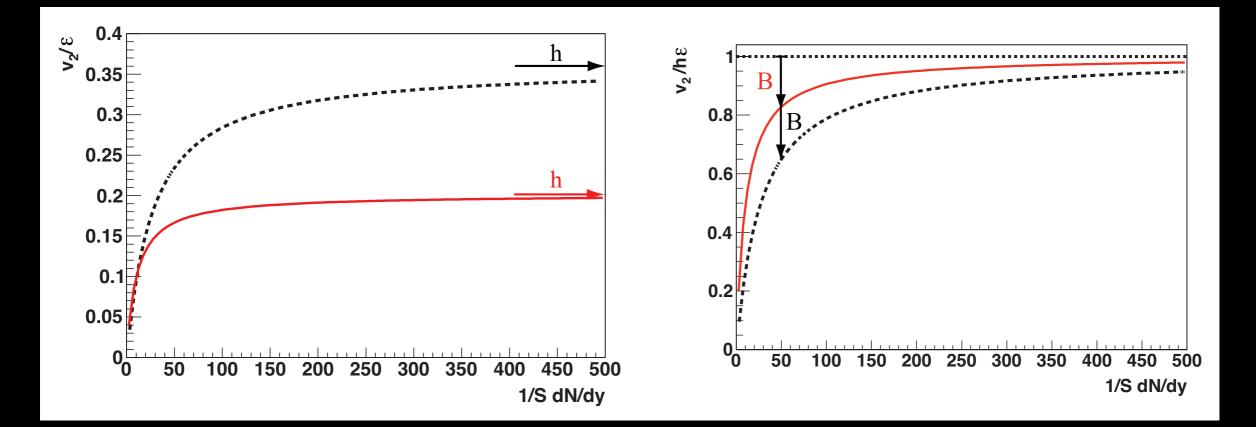
## Integrated V<sub>2</sub>





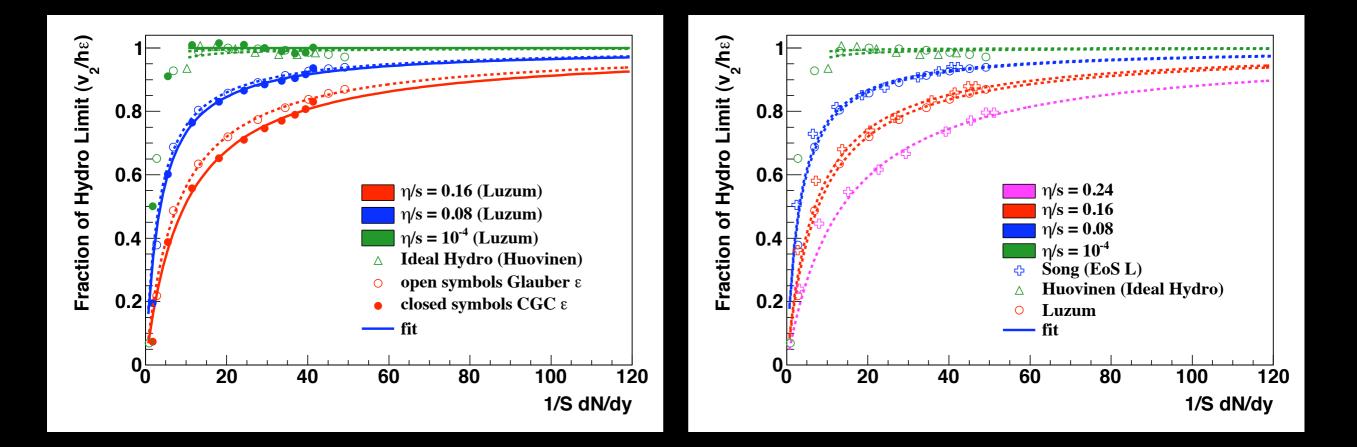


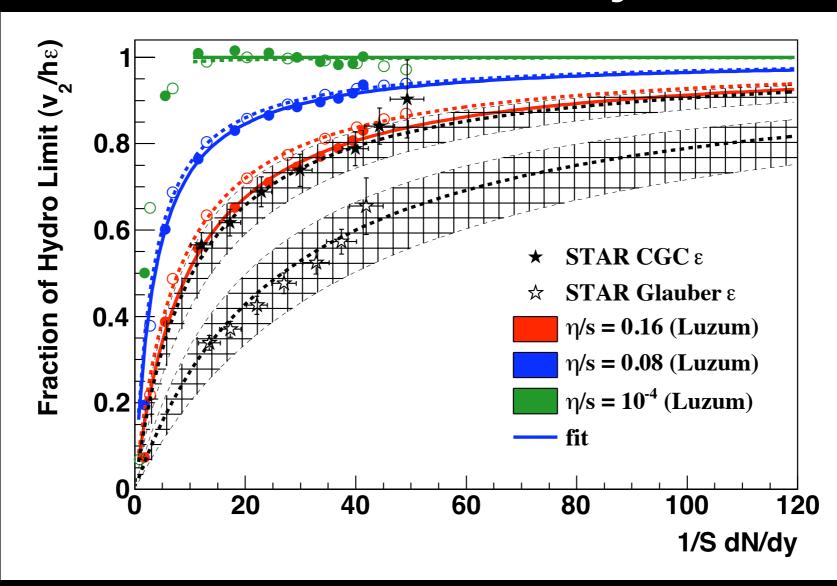
sensitive to the EoS and transport parameters



$$v_n \propto \varepsilon_n$$

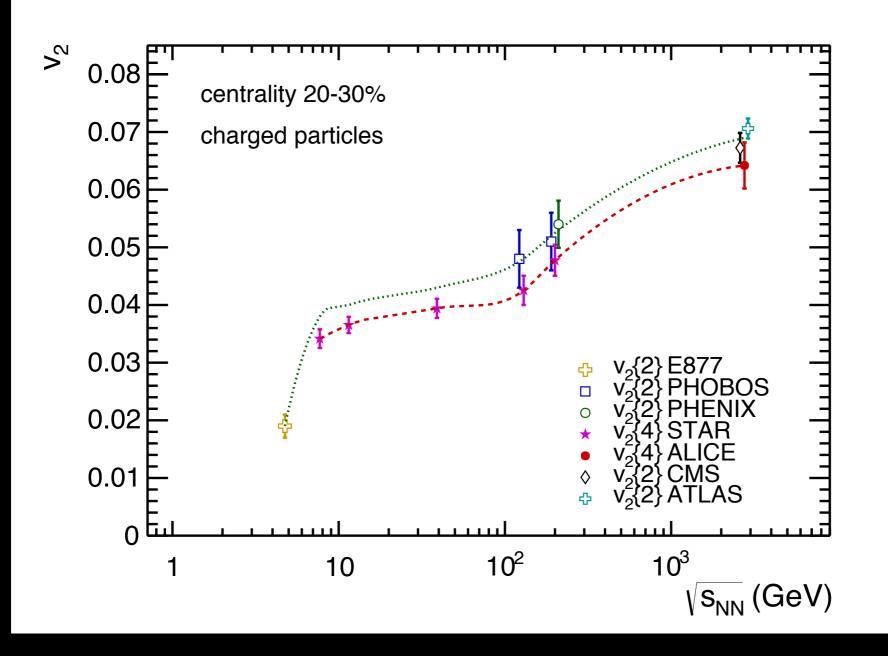
#### sensitive to the EoS and transport parameters





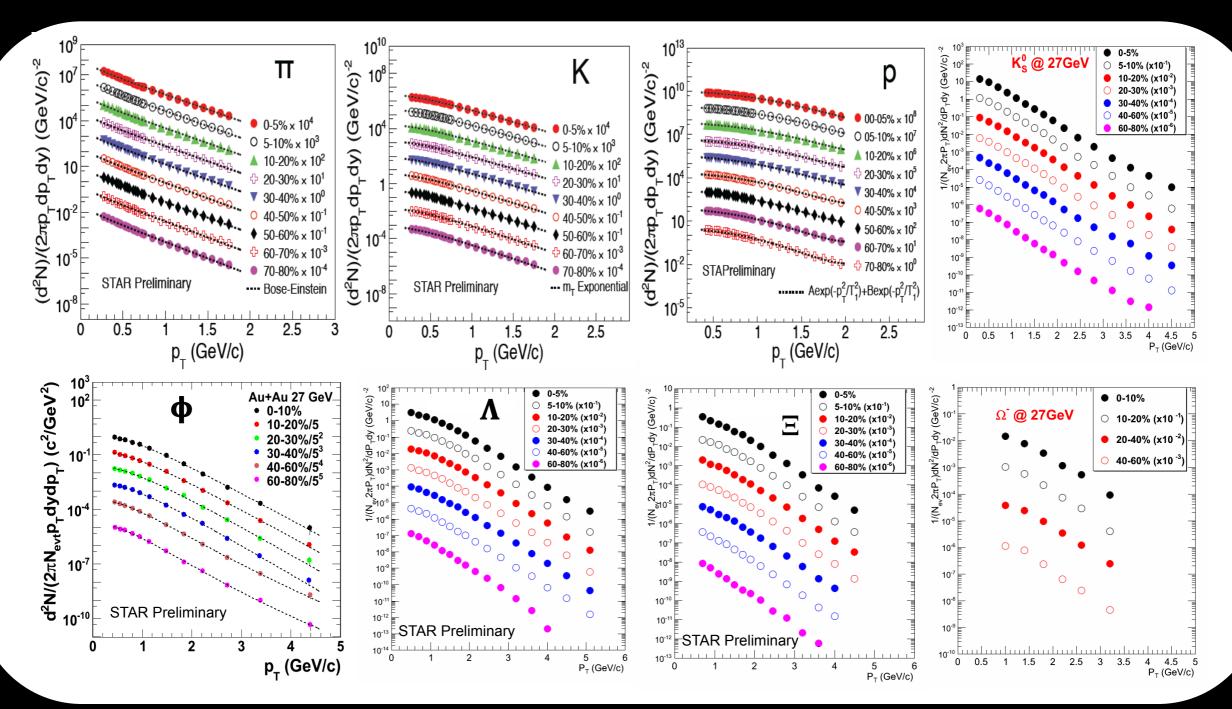
If the models match the data depends strongly on what the true eccentricity is (still an open question)

## Integrated V<sub>2</sub>



collision energy dependence of the elliptic flow shows indication of changing slope

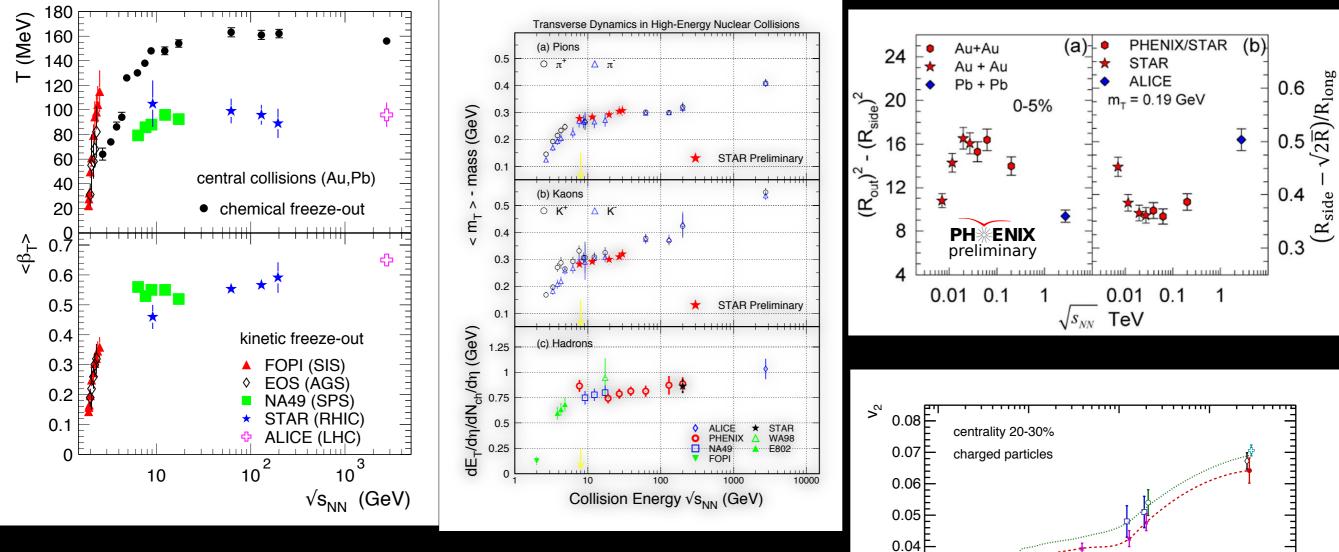
### kinetic freeze-out



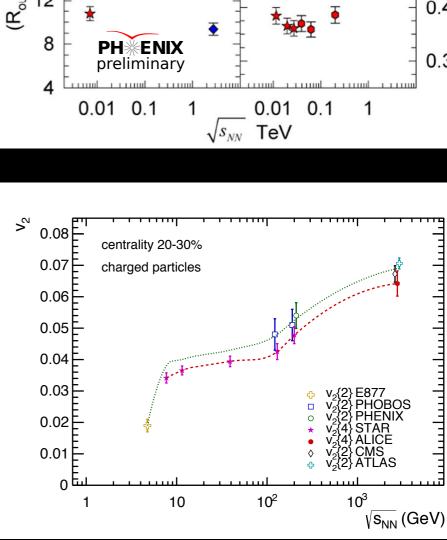
spectra a various particles follow trend expected from a boosted "thermal" system

### kinetic freeze-out

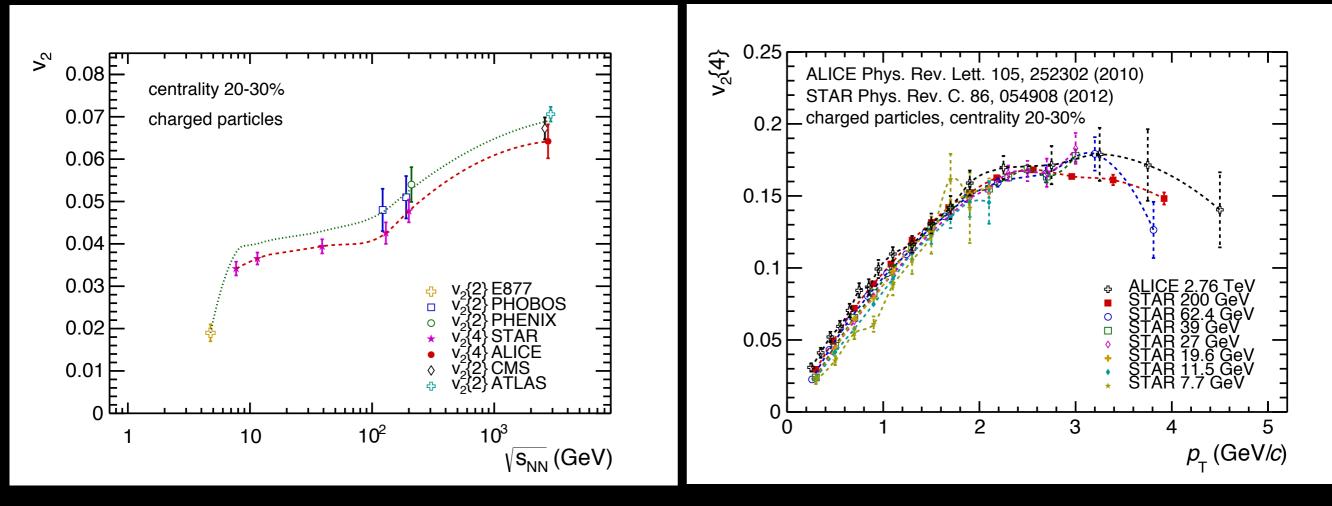
Anton Andronic, Int. J. Mod. Phys. A29 (2014) 1430047



collision energy dependence shows nice indications of changing slope in many observables



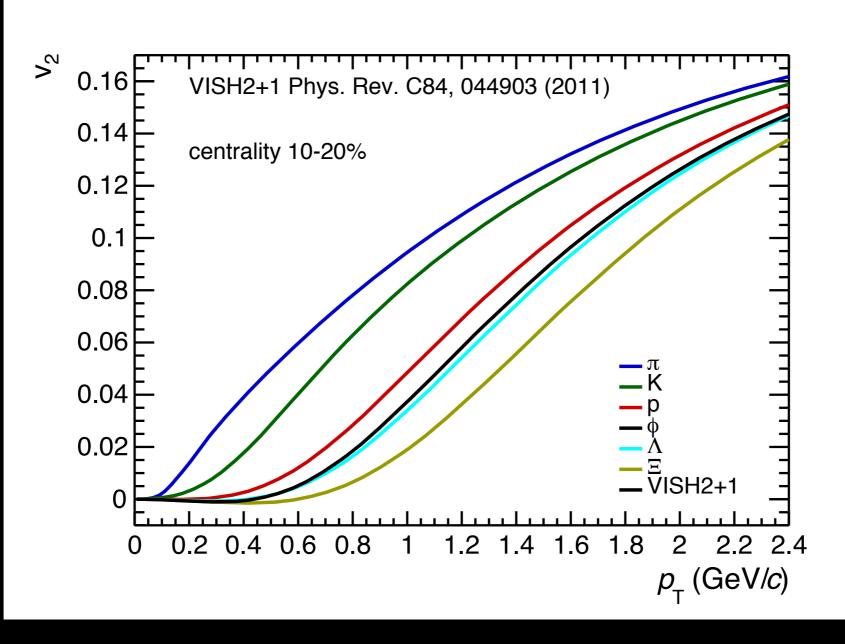
## The increase of collective flow with collision energy?



Elliptic flow increases from RHIC to LHC collision energies about 30% Detailed measurements of v<sub>2</sub>{4} at RHIC in the beam energy scan combined with the LHC measurements show tantalising evidence for a change in slope.

The p<sub>T</sub>-differential elliptic flow also increases with collision energy but difference is small over two orders of magnitude Is this expected/understood?

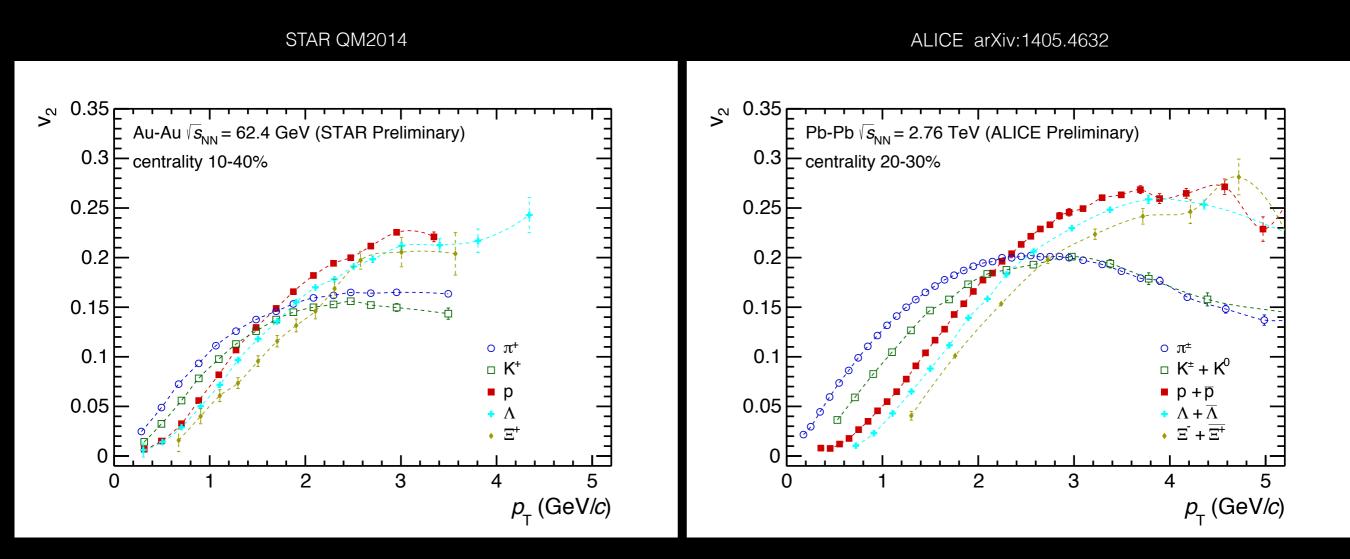
### Collective behaviour



In the hydro and blast-wave picture

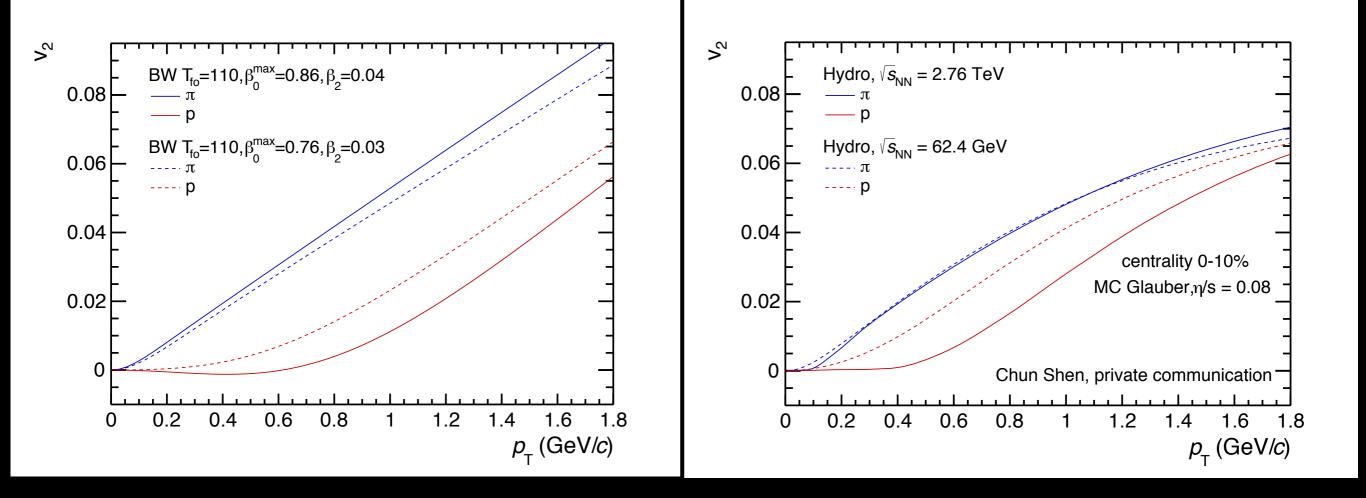
particles have a common temperature and flow velocity at freeze-out. The difference in p⊤-differential elliptic flow depends mainly on one parameter: the mass of the particle

## Collision energy dependence of elliptic flow for particles with different masses



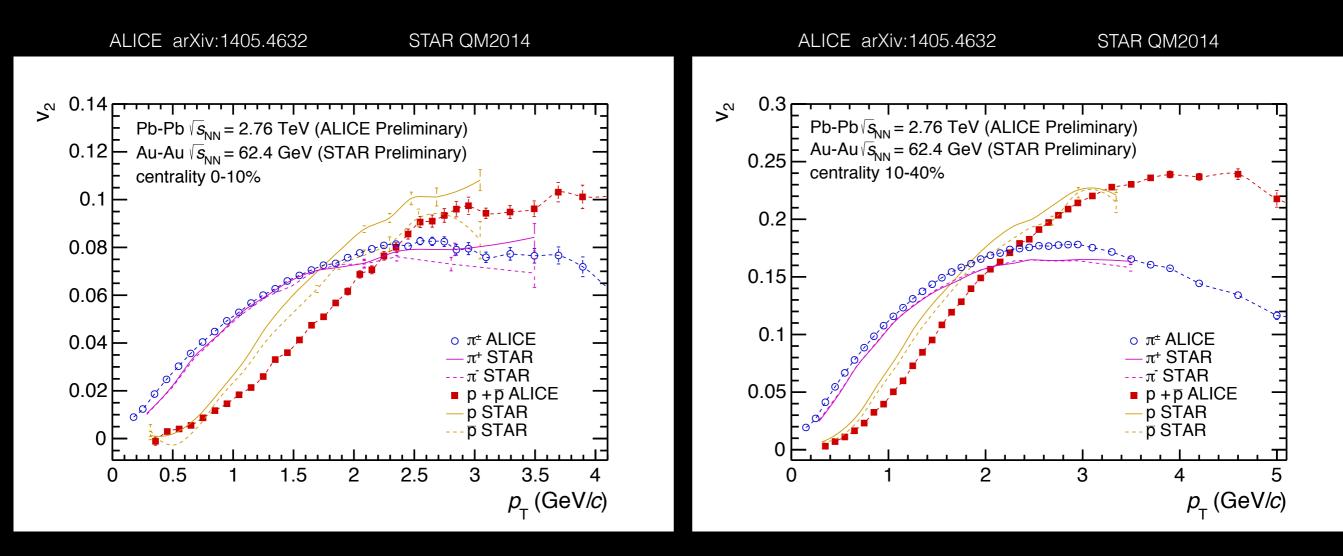
mass hierarchy follows hydrodynamic and blast-wave picture at low p<sub>T</sub>!

### Hydrodynamic behaviour



hydro and blast-wave picture particles have a common temperature and flow velocity larger radial flow increases mass splitting

## Collision energy dependence of elliptic flow as function of transverse momentum

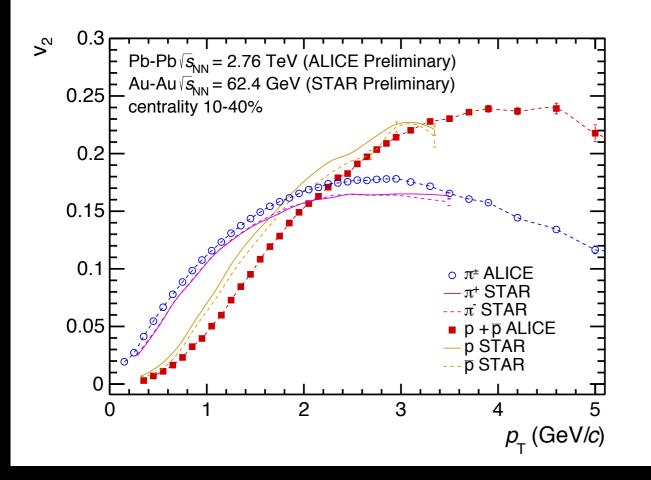


while the  $p_T$ -differential charged particle  $v_2$  changes very little over two orders of magnitude the  $v_2$  of heavier particles clearly shows the effect of the larger collective flow at higher collision energies

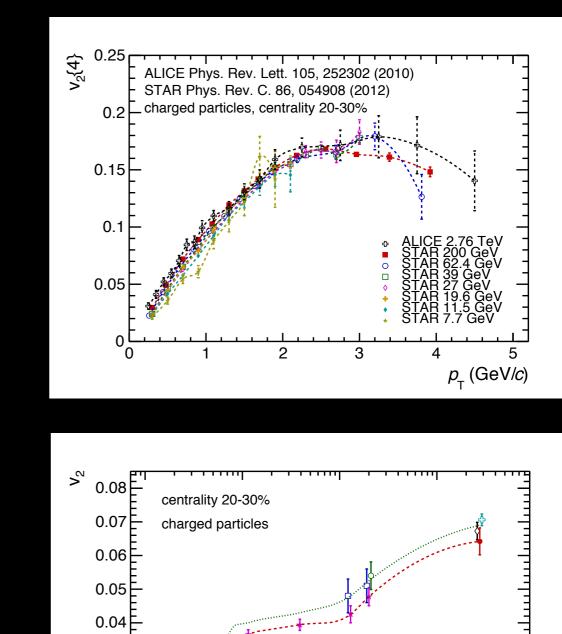
## Collision energy dependence of elliptic flow as function of transverse momentum

ALICE arXiv:1405.4632

STAR QM2014



Elliptic flow as function of collision energy can be qualitatively understood in terms of a boosted thermal system



PHOBOS

'HENIX

 $\sqrt{s_{NN}}$  (GeV)

STAR ALICE

10<sup>3</sup>

 $10^{2}$ 

10



0.03

0.02

0.01

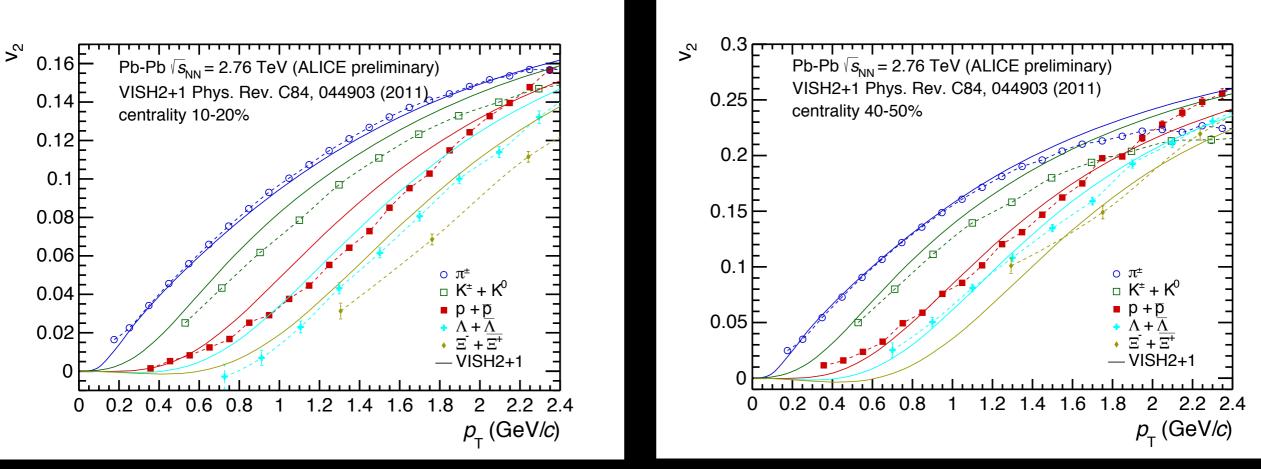
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### Compared to viscous hydrodynamics

ALICE arXiv:1405.4632

ALICE arXiv:1405.4632

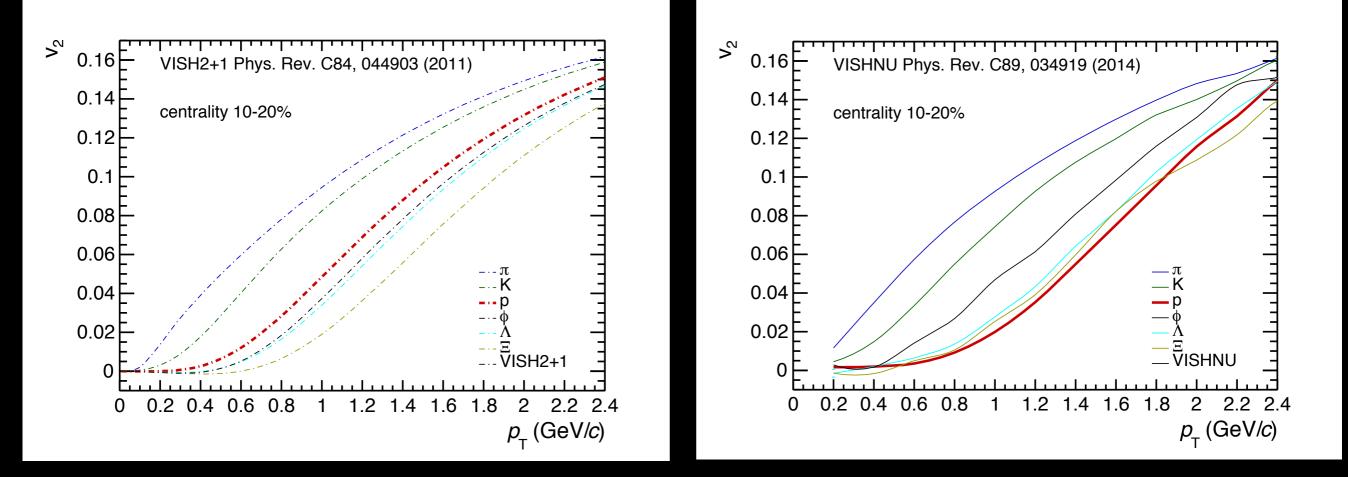


pure viscous hydrodynamics VISH2+1, status at QM2011

Viscous hydrodynamics predictions worked reasonably well for more peripheral collisions 40-50%

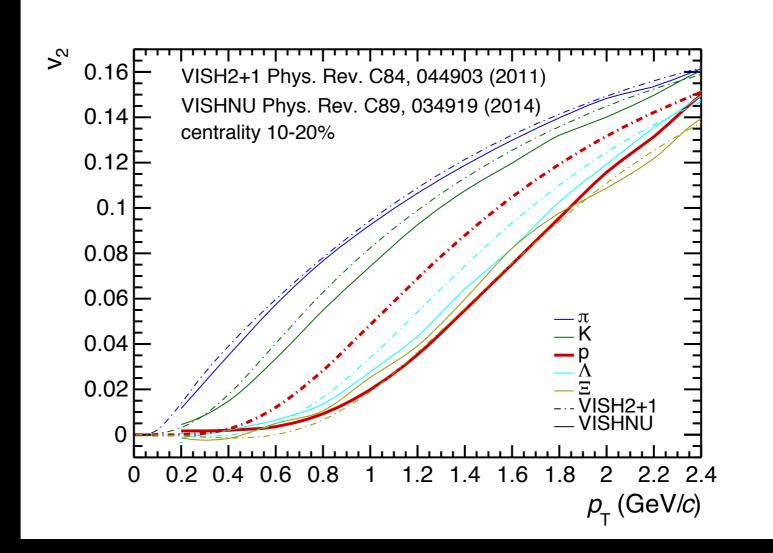
For more central collisions, 10-20%, the radial flow seems to be under-predicted as the protons deviate a lot and this was part of the proton puzzle (the new data plotted here shows this is not just for protons but all heavy particles) can this be understood by a more dissipative hadronic phase (model with a hadron cascade)?

## Viscous hydrodynamics and the effect of the hadronic cascade



VISH2+1 viscous hydrodynamics "standard" mass scaling VISHNU viscous hydrodynamics + hadron cascade mass scaling broken, depending on individual hadronic reinteraction cross sections (pion wind pushing the protons)

## Viscous hydrodynamics and the effect of the hadronic phase



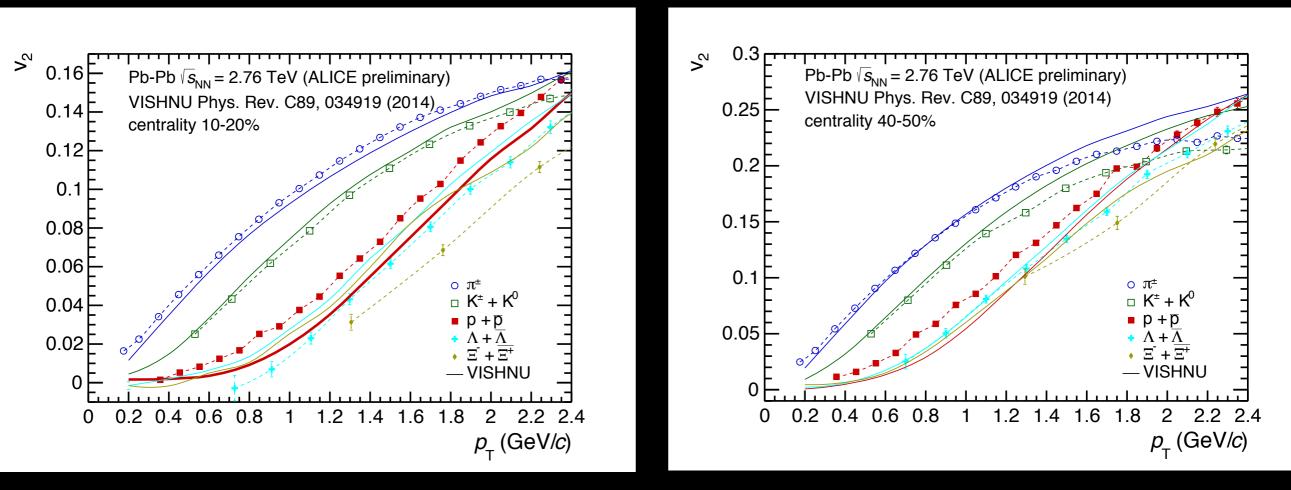
VISHNU viscous hydrodynamics + hadron cascade big effect for the protons! mass scaling broken,

depending on individual hadron-hadron re-interaction cross sections

#### Viscous hydrodynamics + hadron cascade

ALICE arXiv:1405.4632

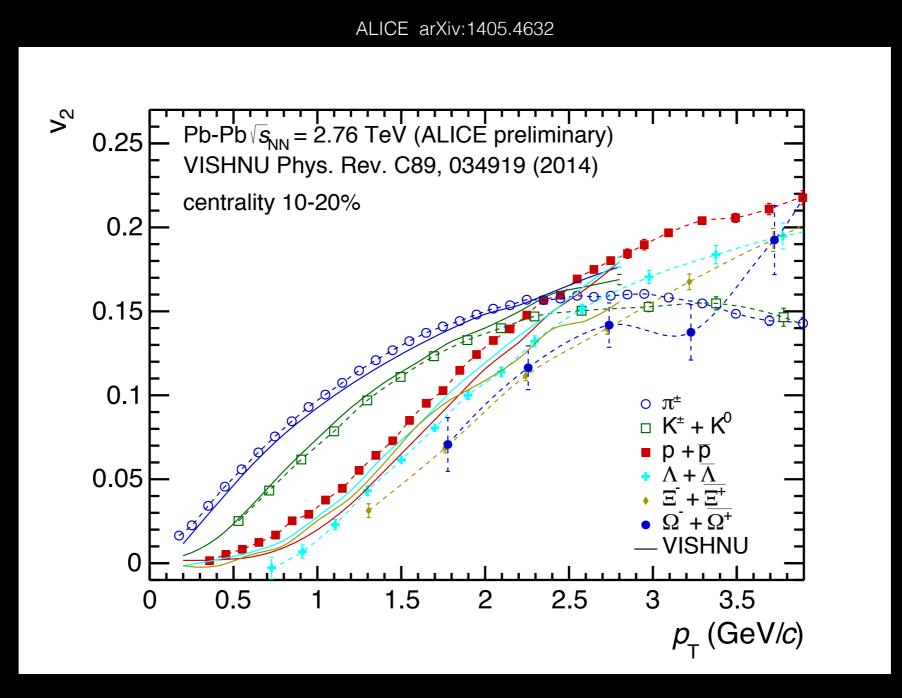
ALICE arXiv:1405.4632



Viscous hydro +hadron cascade improves the Kaon v<sub>2</sub> It increases the push for the protons but actually over does it It breaks the mass scaling and is incompatible with the data It does a worse job than "simple" viscous hydrodynamics!!

over estimating effect of hadronic cascade? or is the model lacking pre-equilibrium flow (AdS/CFT, CGC, .....)?

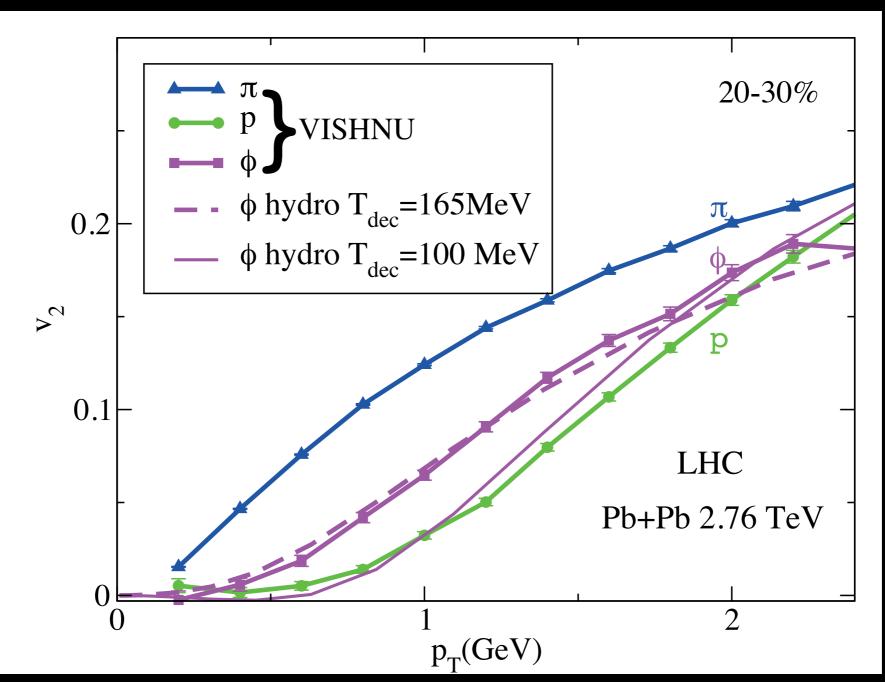
#### Viscous hydrodynamics + hadron cascade



VISHNU does not match with any of the baryons including the  $\Omega$ 's

#### φ-meson elliptic flow

H. Song, S. Bass, U.W. Heinz, Phys. Rev. C89 034919 (2014)

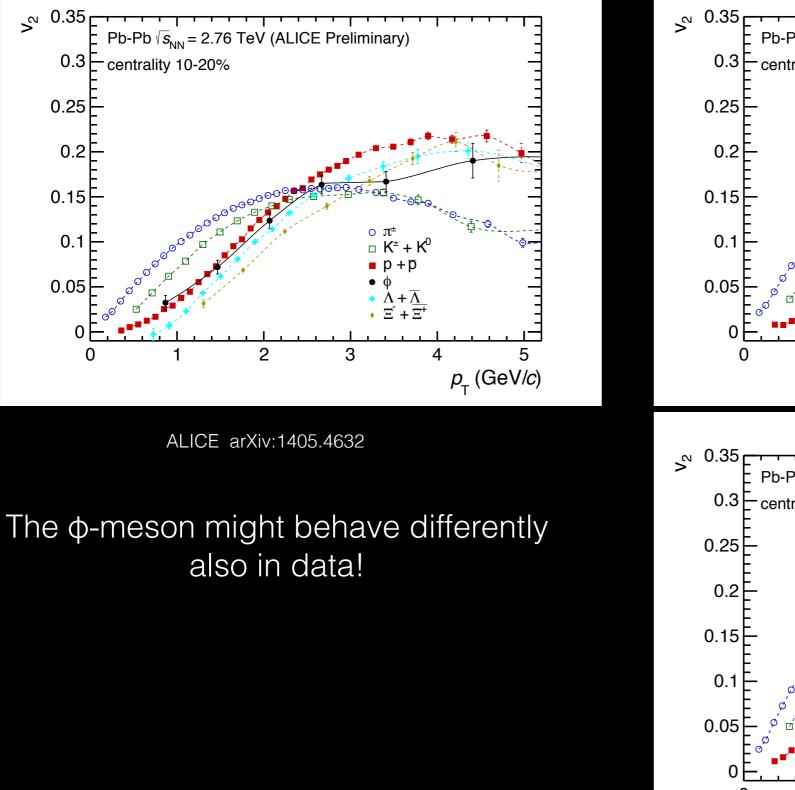


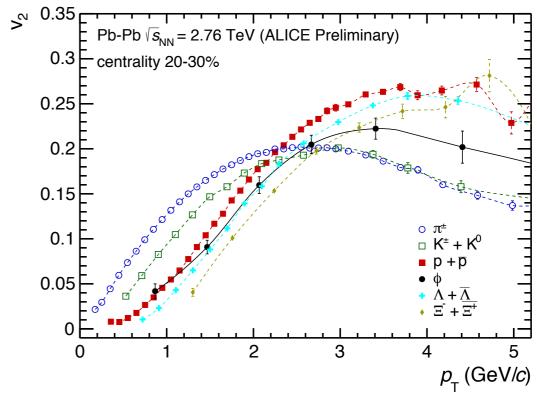
Individual hadronic cross sections matter

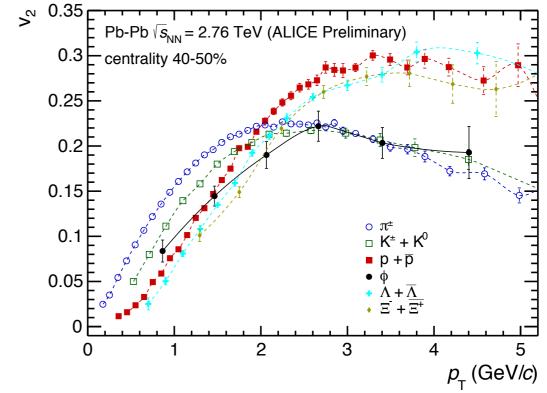
The φ-meson calculations show clear differences (expected because this is put in the model, not a priory clear if this has anything to do with reality)!

### φ-meson elliptic flow

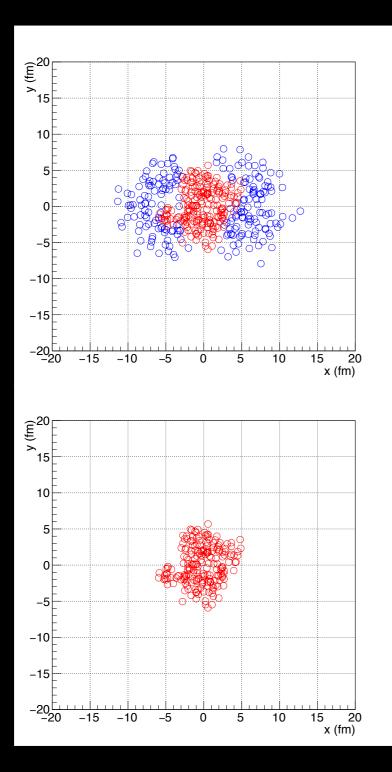
35

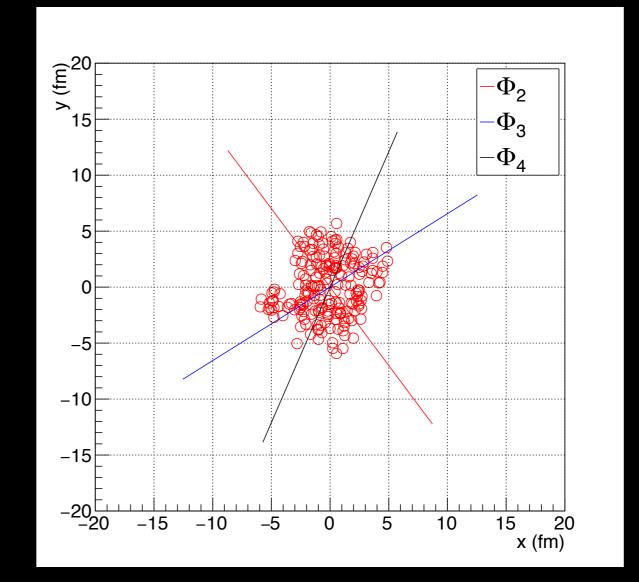






## Symmetry Planes

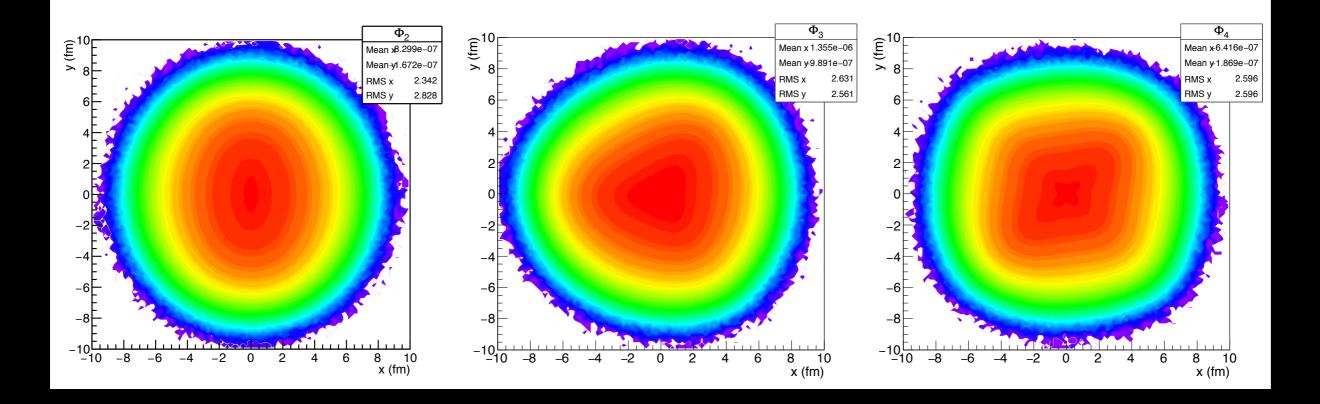




There are many more symmetry planes

 $v_n \propto \varepsilon_n$ 

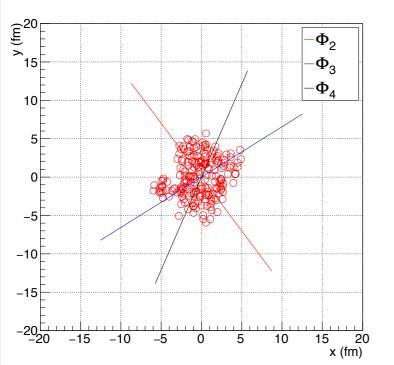
### Symmetry Planes



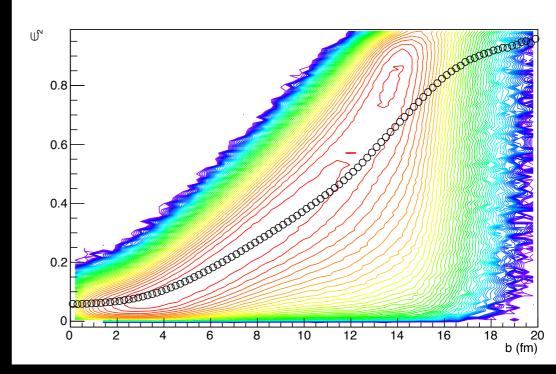
rotated to the planes of symmetry we clearly see the different harmonics

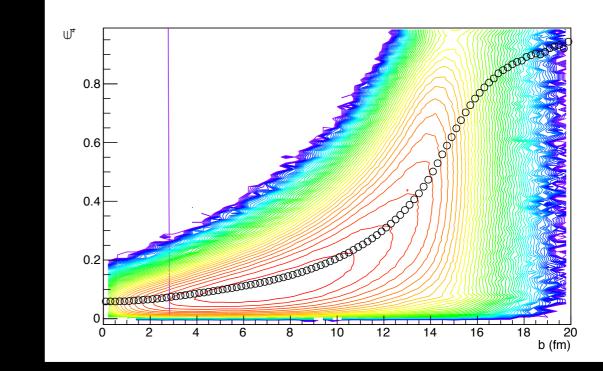
$$v_n \propto \varepsilon_n$$

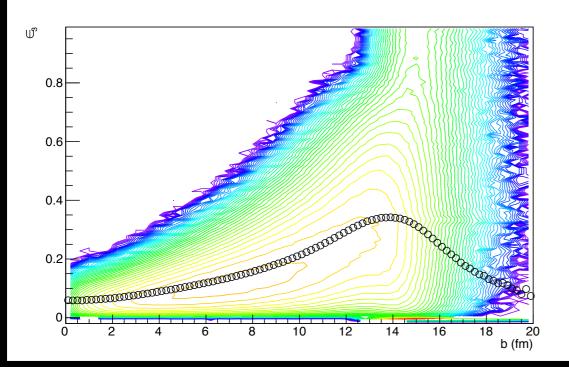
### Eccentricities



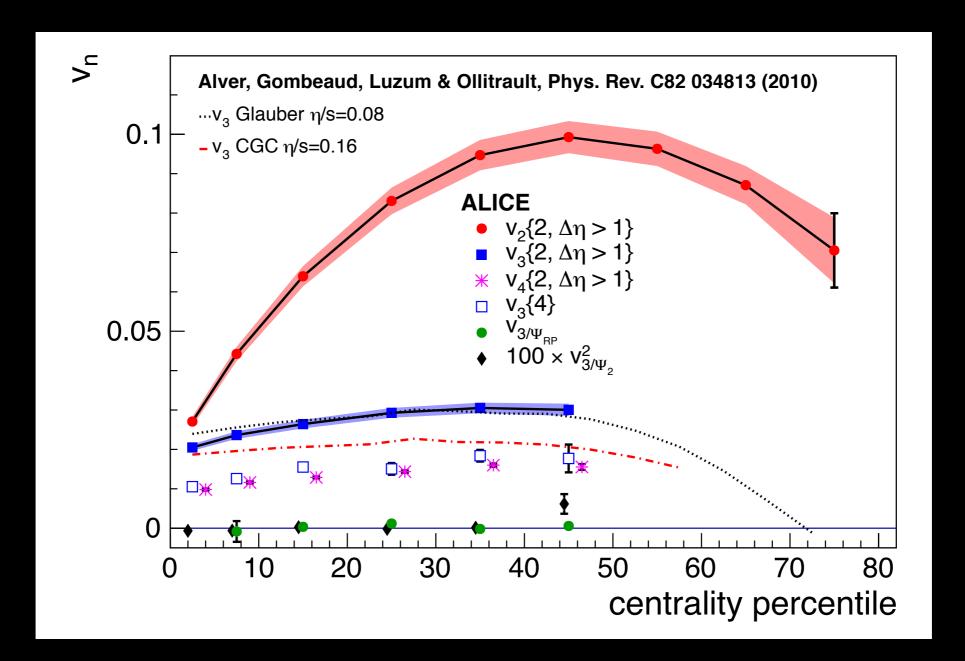




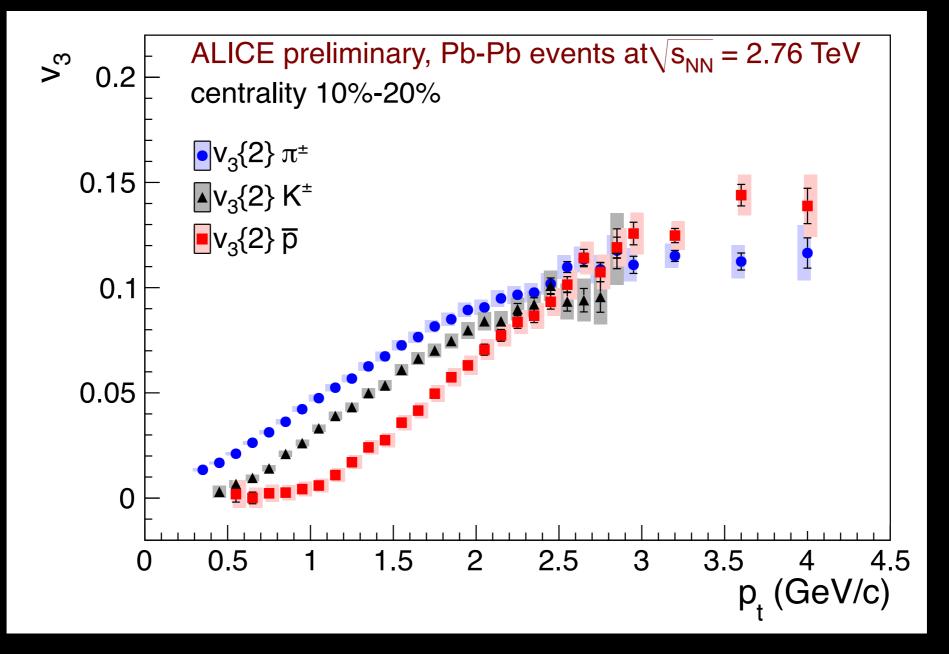




### Higher harmonics



### Higher harmonics



the mass ordering is also observed for higher harmonics

We do not know the reaction plane  $\Psi_R$  or in more general  $\Psi_n$  $v_n \equiv \langle e^{in(\varphi - \Psi_n)} \rangle$ 

We can calculate these observables only using correlations

$$\langle \langle e^{in(\varphi_1 - \varphi_2)} \rangle \rangle = \langle \langle e^{in(\varphi_1)} \rangle \rangle \langle \langle e^{in(\varphi_2)} \rangle \rangle + \langle \langle e^{in(\varphi_1 - \varphi_2)} \rangle_c \rangle$$

zero for symmetric detector when averaged over many events

$$\begin{split} \langle \langle e^{in(\varphi_1 - \varphi_2)} \rangle \rangle &= & \langle \langle e^{in(\varphi_1 - \Psi_n - (\varphi_2 - \Psi_n))} \rangle \rangle \\ &= & \langle \langle e^{in(\varphi_1 - \Psi_n)} \rangle \langle e^{-in(\varphi_2 - \Psi_n)} \rangle \rangle \\ &= & \langle v_n^2 \rangle \end{split}$$

when only  $\Psi_n$  correlations are present

Build cumulants with multi-particle correlations (Ollitrault and Borghini, 2000)

$$c_n\{2\} \equiv \left\langle \left\langle e^{in(\phi_1 - \phi_2)} \right\rangle \right\rangle = v_n^2 + \delta_2$$

$$c_n\{4\} \equiv \left\langle \left\langle e^{in(\phi_1 + \phi_2 - \phi_3 - \phi_4)} \right\rangle \right\rangle - 2 \left\langle \left\langle e^{in(\phi_1 - \phi_2)} \right\rangle \right\rangle^2$$

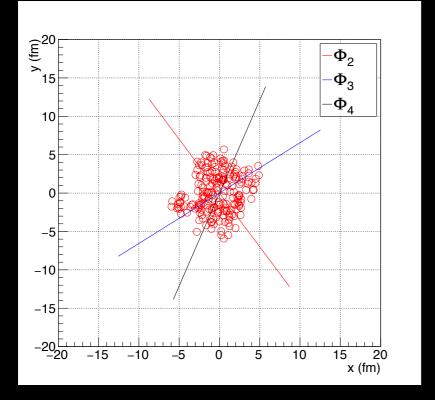
$$= v_n^4 + 4v_n^2 \delta_2 + 2\delta_2^2 - 2(v_n^2 + \delta_2)^2$$

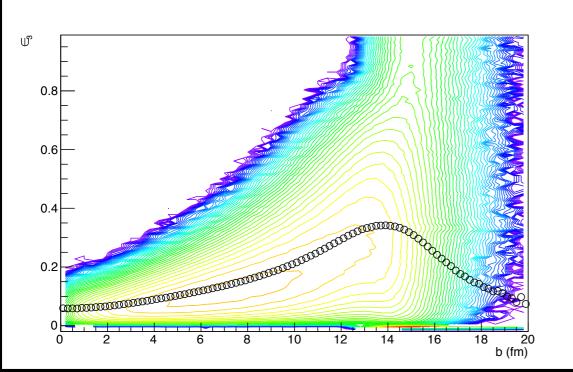
$$= -v_n^4$$

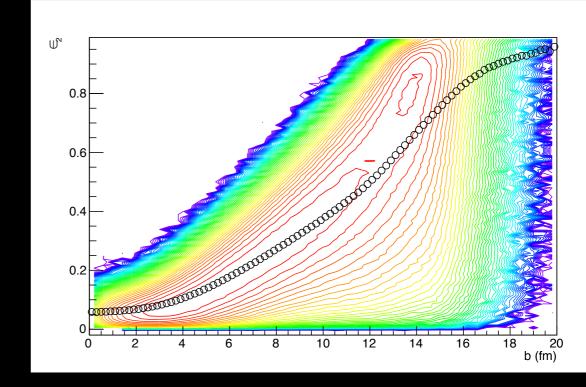
got rid of 2-particle correlations not related to collective flow however now we measure higher moment moments of the distribution

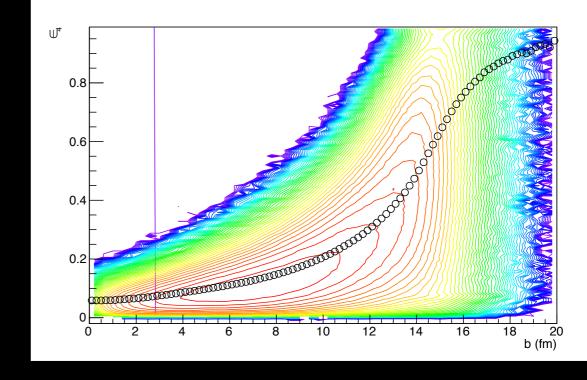
if the fluctuations are small we can say for any distributions that the various flow estimates follow:

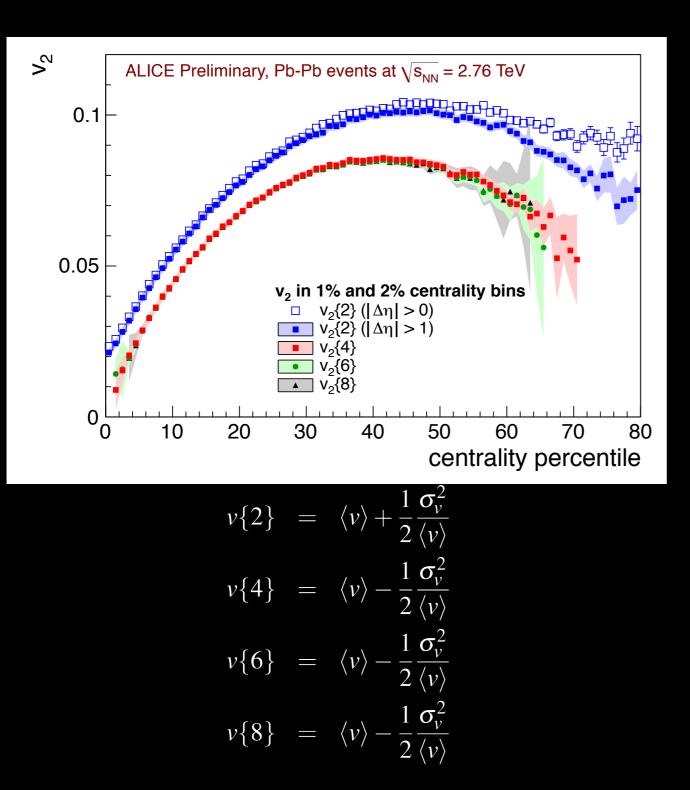
$$v\{2\} = \langle v \rangle + \frac{1}{2} \frac{\sigma_v^2}{\langle v \rangle}$$
$$v\{4\} = \langle v \rangle - \frac{1}{2} \frac{\sigma_v^2}{\langle v \rangle}$$
$$v\{6\} = \langle v \rangle - \frac{1}{2} \frac{\sigma_v^2}{\langle v \rangle}$$
$$v\{8\} = \langle v \rangle - \frac{1}{2} \frac{\sigma_v^2}{\langle v \rangle}$$

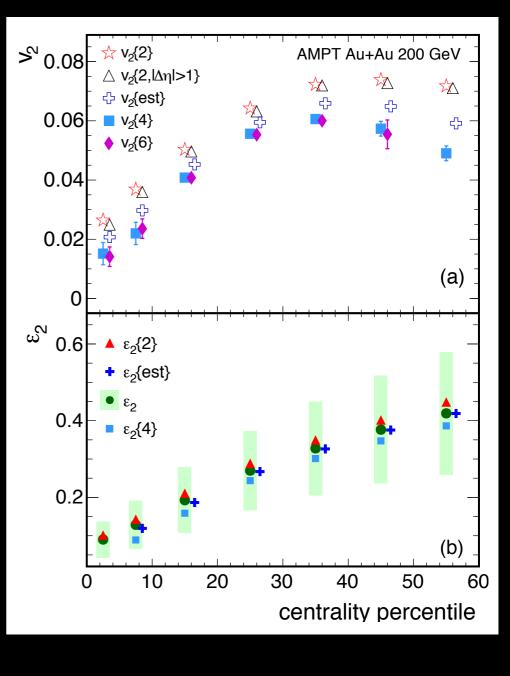


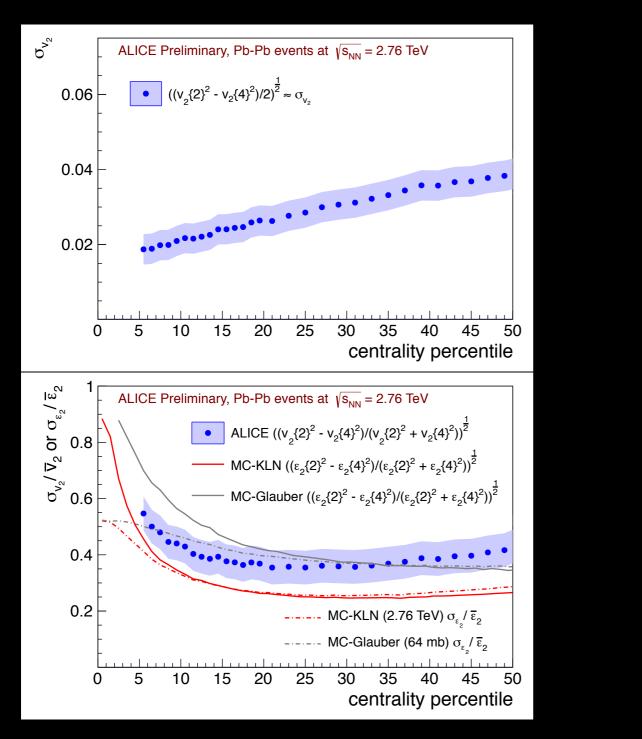


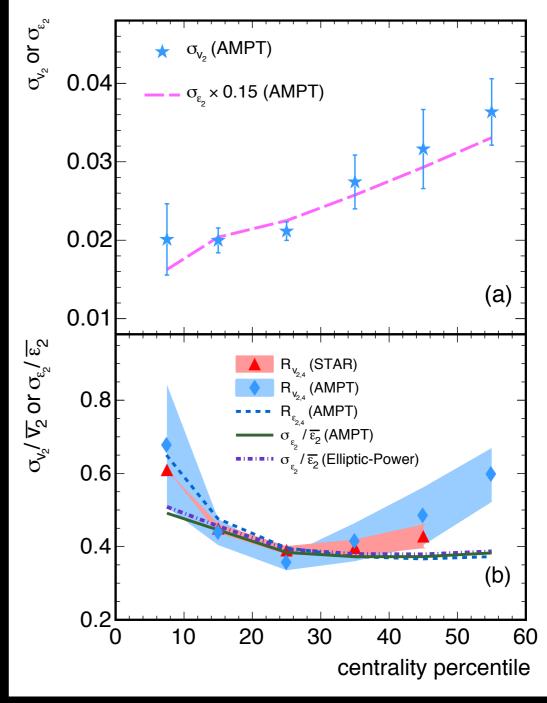












however the fluctuations are rather large

- small fluctuations easy but not much info
- need the underlying pdf!

#### Bessel-Gaussian

$$p(\varepsilon_n) = \frac{\varepsilon_n}{\sigma^2} I_0\left(\frac{\varepsilon_n \varepsilon_n}{\sigma^2}\right) \exp\left(-\frac{\varepsilon_0^2 + \varepsilon_n^2}{2\sigma^2}\right)$$

 ε<sub>0</sub> is the anisotropy versus the reaction plane and σ the fluctuations.
 Works for mid-central collisions, not
 expected to work for peripheral collisions because not constraint to 1 this distribution predict that v<sub>3</sub>{4}=0

Power-law distribution

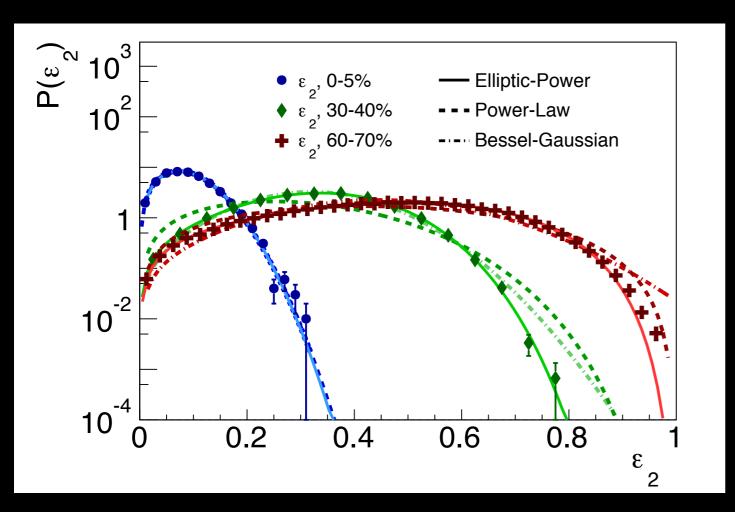
 $p(\varepsilon_n) = 2\alpha \,\varepsilon_n \left(1 - \varepsilon_n^2\right)^{\alpha - 1}$ 

 $\alpha$  quantifies the fluctuations, this function has no  $\epsilon_0$  therefore only describes the response due to fluctuations

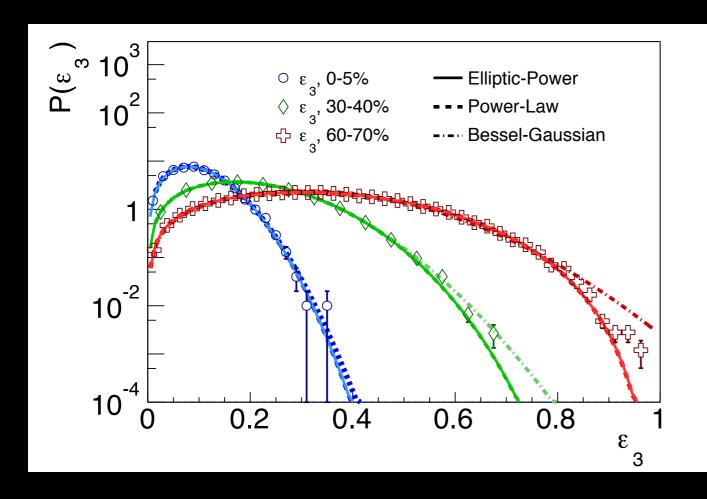
Elliptic Power distribution

$$p(\varepsilon_n) = \frac{\alpha \varepsilon_n}{\pi} \left(1 - \varepsilon_0^2\right)^{\alpha + \frac{1}{2}} \int_0^{2\pi} \frac{\left(1 - \varepsilon_n^2\right)^{\alpha - 1} d\phi}{\left(1 - \varepsilon_0 \varepsilon_n \cos\phi\right)^{2\alpha + 1}}$$

 $\alpha$  and  $\epsilon_0$  are the ingredients with same definition as in previous distributions

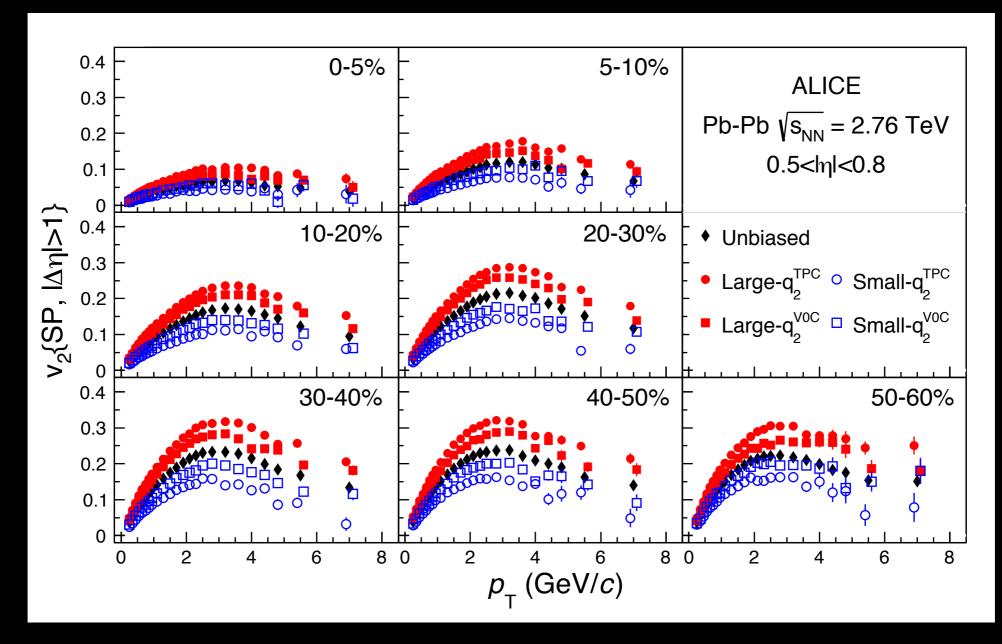


In 0-5% all three functions work rather well. This is understood,  $\varepsilon_0$  is small and a is large. Elliptic Power turns into a Bessel Gaussian and with  $\varepsilon_0$  small the anisotropy versus the reaction plane and power law also works. For more peripheral collisions the Elliptic Power is the only distributions which works well



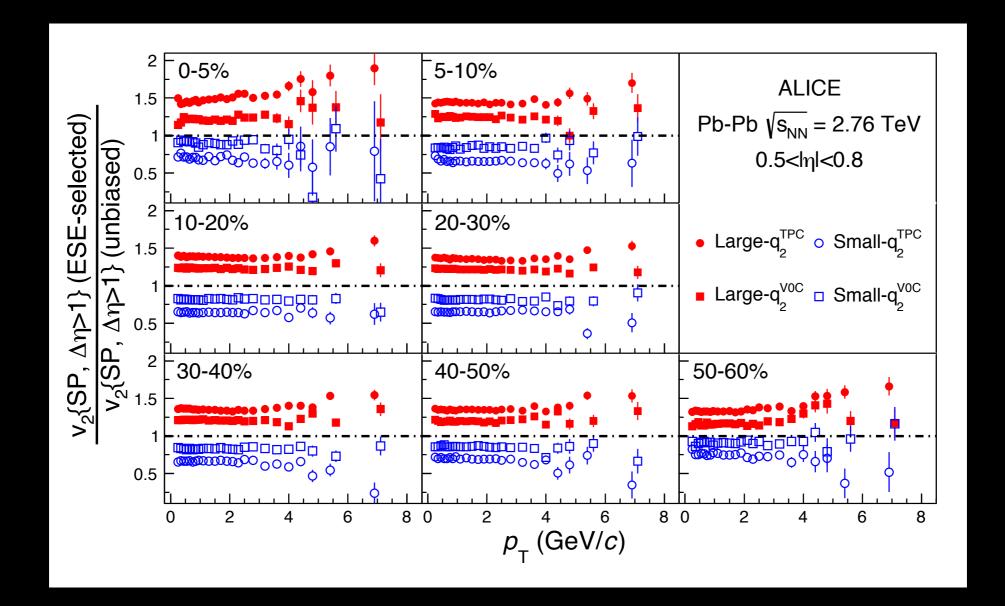
 $\epsilon_3$ ,  $v_3$  dominated by fluctuations. For more central collisions all three functions work rather well. Again this is understood Bessel Gaussian fails for more peripheral due to lack of constraint < 1. The fact that  $\epsilon_3$ {4} and  $v_3$ {4} are non-zero completely excluded the Bessel Gaussian

### Event shape engineering

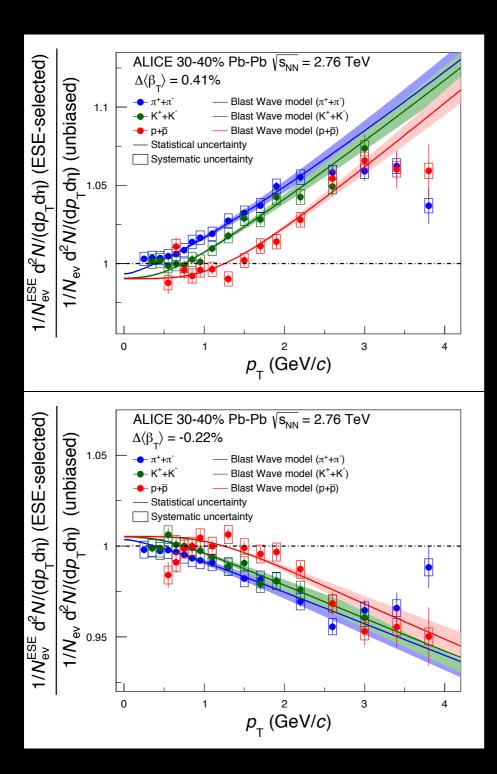


Can use the fluctuations to select events with different shapes at fixed centralities Perfect tool to test the response of the system outside of our normal selection on centrality and/or collision system

### Event shape engineering



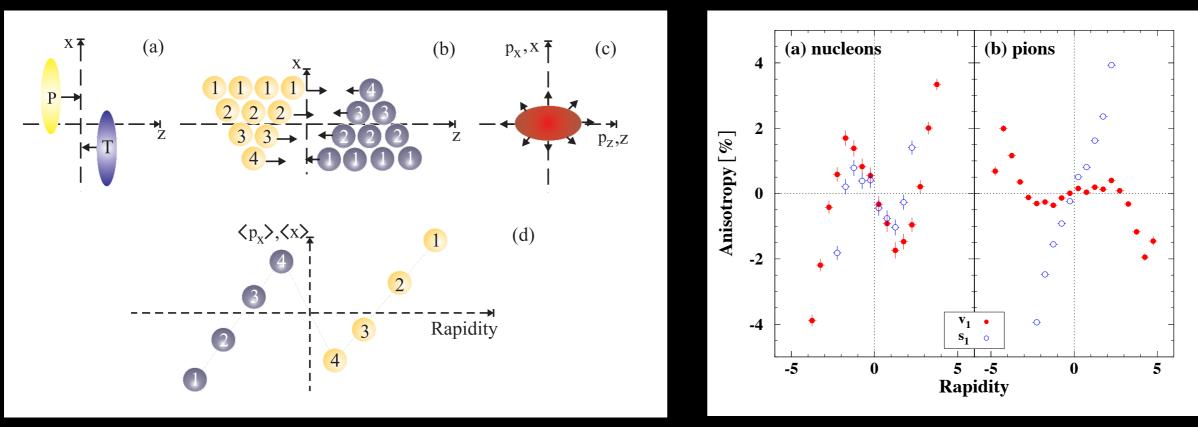
### Event shape engineering



The radial flow indeed also changes for "identical" collisions which only differ in shape new and starting field!

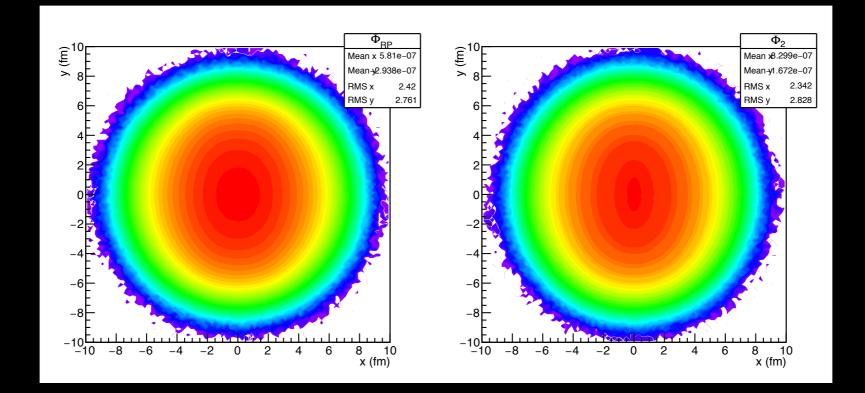
### Rapidity Dependence

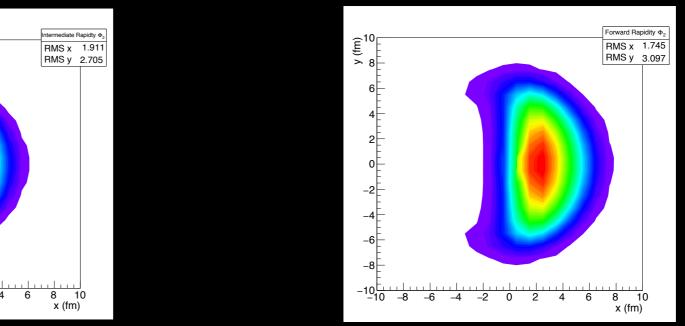
R.S, H. Sorge, S.A. Voloshin, F.Q. Wang, N. XU; PRL 84 (2000)

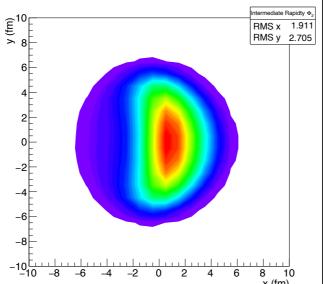


#### stopping can be important for how the initial spatial distributions looks like

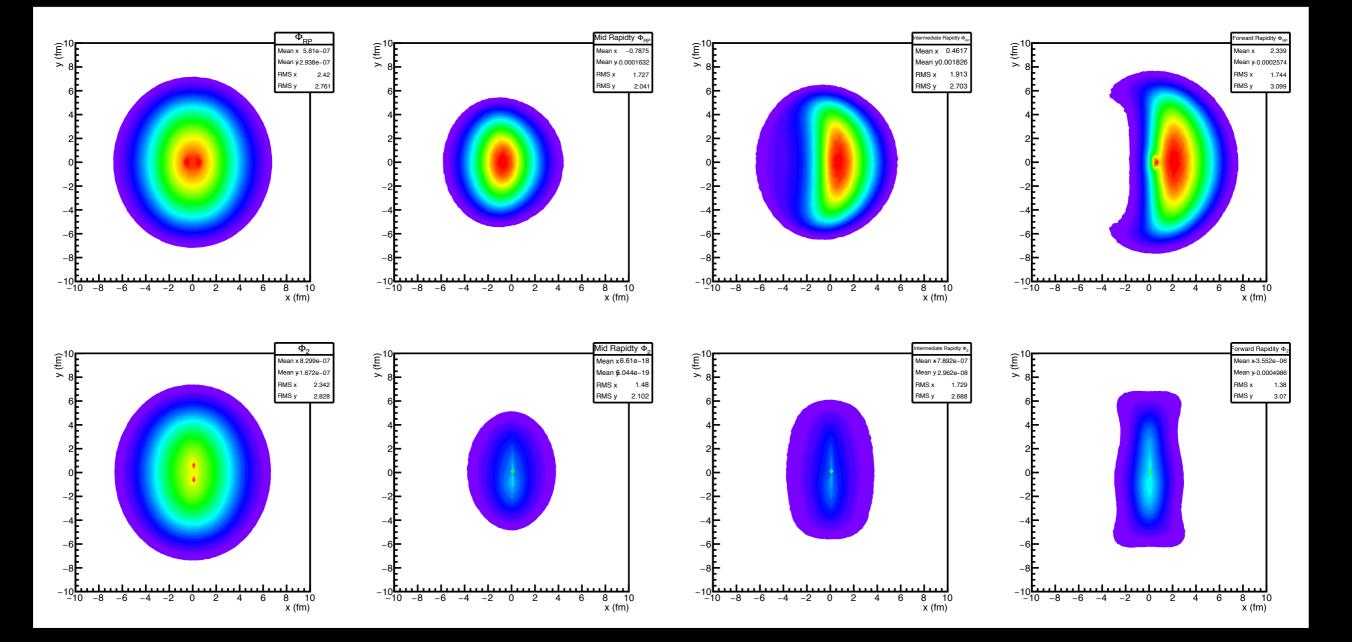
### Rapidity Dependence







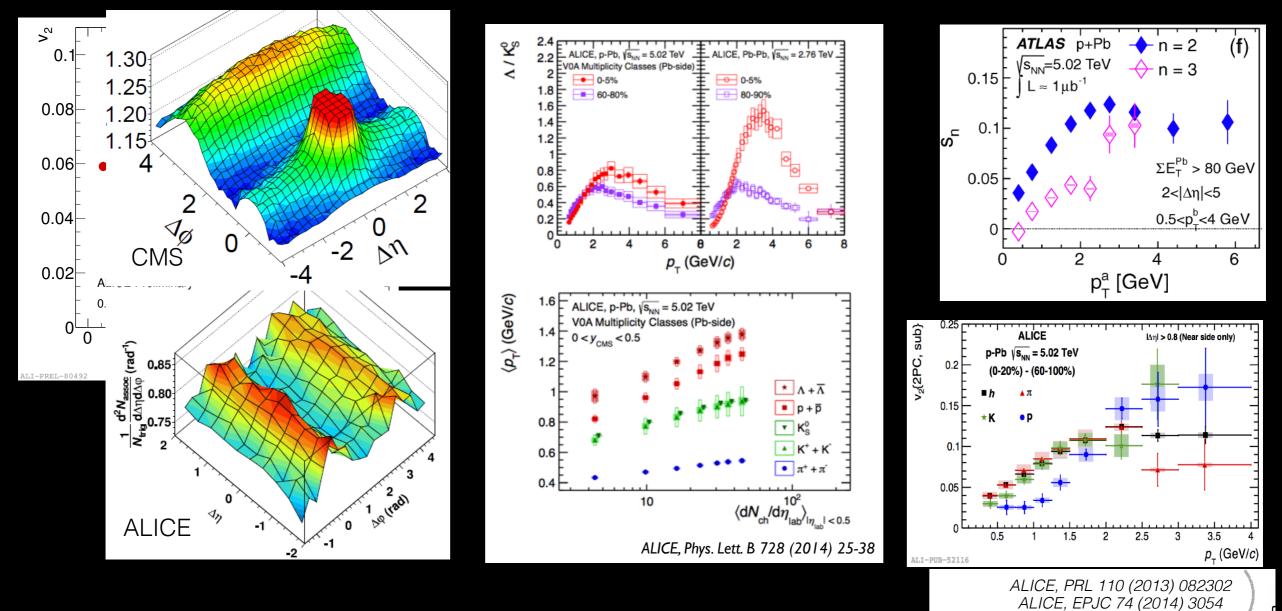
## Rapidity Dependence



### Small systems; pA collisions

- a reference for AA (cold nuclear matter effects)
- a ideal system for CGC studies

### pA collisions



- collective effects in pA?
  - who ordered that?

 $R_{\mathsf{pPb}}$ 

**1.3**₿

1.2 1.1

0.9 0.8 0.7

0

5

ALICE p-Pb  $\sqrt{s_{NN}}$ =5.02 TeV, NSD

 $|\eta_{mc}| < 0.3$ 

20

25

30

35

15

10

p-Pb

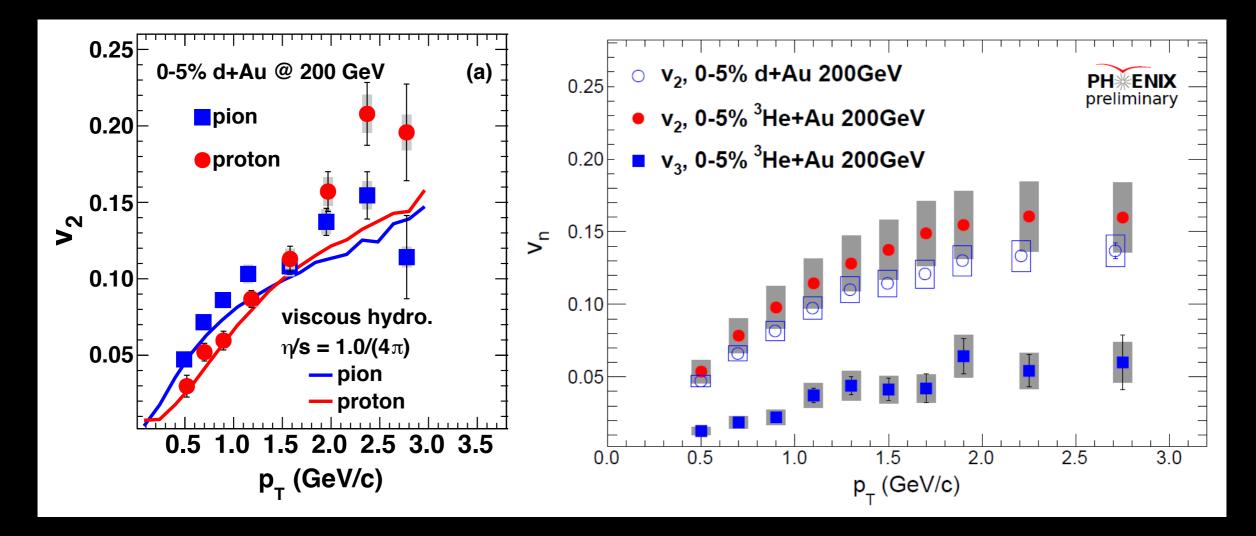
45

 $p_{_{\rm T}}$  (GeV/c)

40

charged particles

### d+A and <sup>3</sup>He+A collisions



- collective effects in dA and <sup>3</sup>He+A at RHIC?
  - who ordered that?

### a rose?

- it smells like a rose
- it pricks like a rose







### Collective motion

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#### Generalized Cumulant Expansion Method\*

Ryogo Kubo

Department of Physics, University of Tokyo (Received April 11, 1962)

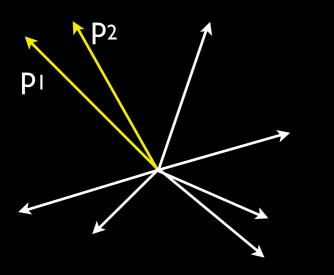
The moment generating function of a set of stochastic variables defines the cumulants or the semi-invariants and the cumulant function. It is possible, simply by formal properties of exponential functions, to generalize to a great extent the concepts of cumulants and cumulant function. The stochastic variables to be considered need not be ordinary c-numbers but they may be q-numbers such as used in quantum mechanics. The exponential function which defines a moment generating function may be any kind of generalized exponential, for example an ordered exponential with a certain prescription for ordering q-number variables. The definition of average may be greatly generalized as far as the condition is fulfilled that the average of unity is unity. After statements of a few basic theorems these generalizations are discussed here with certain examples of application. This generalized cumulant expansion provides us with a point of view from which many existent methods in quantum mechanics and statistical mechanics can be unified.

 $\langle X_j \rangle_{\rm c} = \langle X_j \rangle$  $\langle X_j^2 \rangle_{\rm c} = \langle X_j^2 \rangle - \langle X_j \rangle^2$  $\langle X_j X_l \rangle_c = \langle X_j X_l \rangle - \langle X_j \rangle \langle X_l \rangle$  $\langle X_j X_k X_l \rangle_{\rm c} = \langle X_j X_k X_l \rangle$  $-\{\langle X_j \rangle \langle X_k X_l \rangle + \langle X_k \rangle \langle X_l X_j \rangle + \langle X_l \rangle \langle X_j X_k \rangle \}$  $+2\langle X_j \rangle \langle X_k \rangle \langle X_l \rangle$  $\langle X_j X_k X_l X_m \rangle_c = \langle X_j X_k X_l X_m \rangle$  $-\{\langle X_j \rangle \langle X_k X_l X_m \rangle + \langle X_k \rangle \langle X_j X_l X_m \rangle + \langle X_l \rangle \langle X_j X_k X_m \rangle + \langle X_m \rangle \langle X_j X_k X_l \rangle\}$  $-\{\langle X_j X_k \rangle \langle X_l X_m \rangle + \langle X_j X_l \rangle \langle X_k X_m \rangle + \langle X_j X_m \rangle \langle X_k X_l \rangle\}$  $+2\langle\langle X_{j}\rangle\langle X_{k}\rangle\langle X_{l}X_{m}\rangle+\langle X_{j}\rangle\langle X_{l}\rangle\langle X_{k}X_{m}\rangle+\langle X_{j}\rangle\langle X_{m}\rangle\langle X_{k}X_{l}\rangle$  $+\langle X_{j}X_{k}\rangle\langle X_{l}\rangle\langle X_{m}\rangle+\langle X_{j}X_{l}\rangle\langle X_{k}\rangle\langle X_{m}\rangle+\langle X_{j}X_{m}\rangle\langle X_{k}\rangle\langle X_{l}\rangle\}$  $-6\langle X_j \rangle \langle X_k \rangle \langle X_l \rangle \langle X_m \rangle$ 

(2.8)

cumulants allow us to see if there are multi-particle correlations in the system (cumulants nonzero only mathematical proof)

Collective motion  $\langle\!\langle e^{in(\phi_1 - \phi_2)} \rangle\!\rangle = \overline{\langle v_n^2 \rangle} + \delta_2$ 



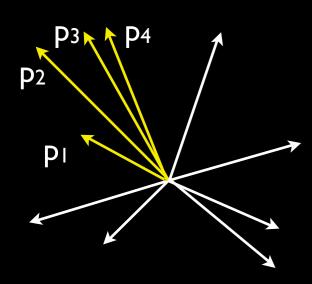
particle I coming from the resonance. Out of remaining M-I particles there is only one which is coming from the same resonance, particle 2. Hence a probability that out of M particles we will select two coming from the same resonance is ~ I/(M-I). From this we can draw a conclusion that for large multiplicity:  $\delta_2 \sim 1/M$ 

• therefore to reliably measure flow:

 $v_n^2 \gg 1/M \Rightarrow v_n \gg 1/M^{1/2}$ 

• not easily satisfied:  $M=200 v_n >> 0.07$ 

### Collective motion



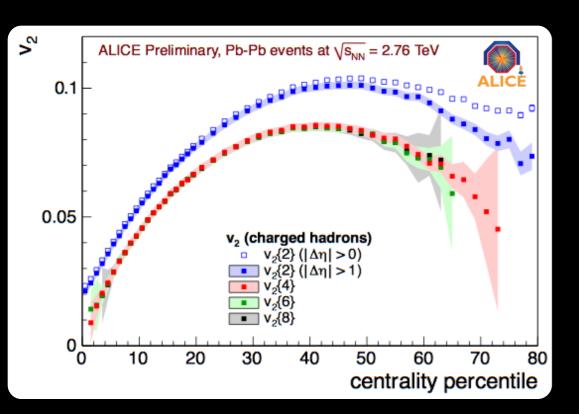
Particle I coming from the mini-jet. To select particle 2 we can make a choice out of remaining M-I particles; once particle 2 is selected we can select particle 3 out of remaining M-2 particles and finally we can select particle 4 out of remaining M-3 particles. Hence the probability that we will select randomly four particles coming from the same resonance is I/(M-I)(M-2) (M-3). From this we can draw a conclusion that for large multiplicity:  $\delta_2 \sim 1/M$ ,  $\delta_4 \sim 1/M^3$ 

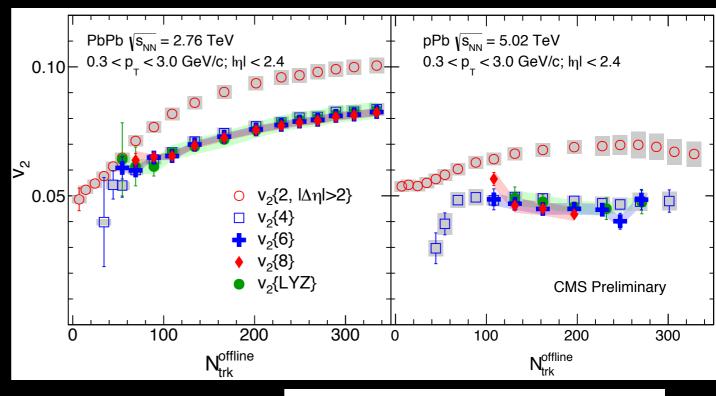
• therefore to reliably measure flow:

$$v_n^2 \gg 1/M \implies v_n \gg 1/M^{1/2}$$
  
 $v_n^4 \gg 1/M^3 \implies v_n \gg 1/M^{3/4}$ 

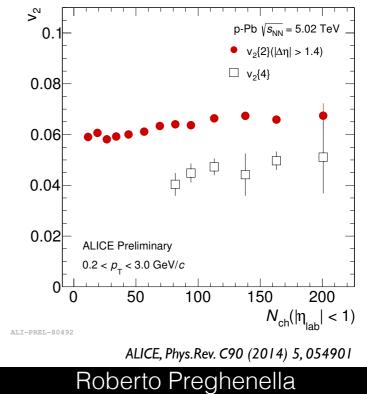
### Collective motion

Stefan Bass



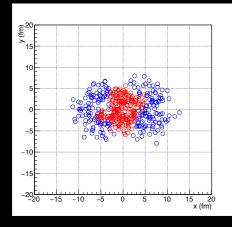


- collective behaviour in pA!
- not necessarily hydrodynamics!
  - quantitative question
- initial state or final state?

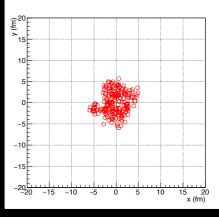


### Who ordered that?

- What can we learn from pA?
  - better theoretical understanding initial state pA compared to AA
  - better experimental constraints on the initial *geometry* in AA compared to pA
  - if both initial state and the final state interactions are important in pA is there still a clear preference compared to peripheral AA?
    - new questions to answer
  - d+A and <sup>3</sup>He+A are very important to disentangle initial or final state origin!
  - (multi-particle) azimuthal correlations differentially as function of  $p_T$  and  $\eta$  should allow us to also test if there are different regimes where initial state or final state effects dominate



# Summary



- clear evidence of the importance of the initial spatial distribution (in all gory details) in all the correlations
- naturally explained if the constituents have strong final state interactions
  - some depend non-trivially on the evolution (which is well captures in models with final state interactions)
- very rich playground for theorist and experimentalist!

# Anisotropic Flow; ad infinitum

**Raimond Snellings** 



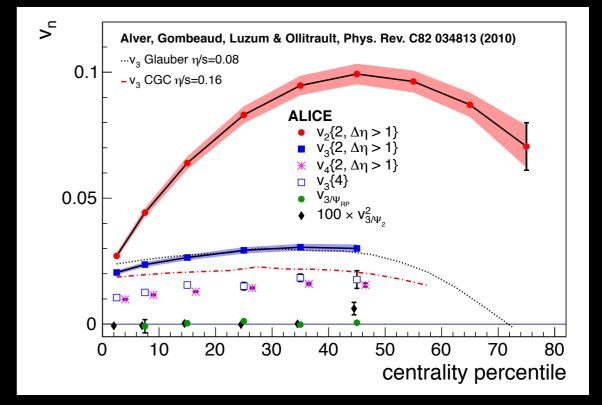
NI<mark>KH</mark>EF

4<sup>th</sup> International Symposium on Non-equilibrium Dynamics 30-08 — 05-09-2015 Giardini Naxos, Sicily,Italy

### Mixed Harmonics and Standard Candles

$$v_{3}\{\Psi_{\rm RP}\} = \langle \langle \cos 3(\varphi - \Psi_{\rm RP}) \rangle \rangle$$
$$= \langle \langle \cos 3(\varphi - \Psi_{3}) \cos 3(\Psi_{3} - \Psi_{\rm RP}) \rangle \rangle$$
$$= \langle v_{3} \langle \cos 3(\Psi_{3} - \Psi_{\rm RP}) \rangle \rangle$$

$$v_3^2\{\Psi_2\} = \frac{\langle \cos(2\varphi_1 + 2\varphi_2 + 2\varphi_3 - 3\varphi_4 - 3\varphi_5) \rangle}{v_2^3}$$



### Mixed Harmonics and Standard Candles

$$v_{3}\{\Psi_{\rm RP}\} = \langle \langle \cos 3(\varphi - \Psi_{\rm RP}) \rangle \rangle$$
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$$= \langle v_{3} \langle \cos 3(\Psi_{3} - \Psi_{\rm RP}) \rangle \rangle$$

$$v_3^2\{\Psi_2\} = \frac{\langle \cos(2\varphi_1 + 2\varphi_2 + 2\varphi_3 - 3\varphi_4 - 3\varphi_5) \rangle}{v_2^3}$$

$$\langle \cos(n_1\varphi_1 + \dots + n_k\varphi_k) \rangle$$
$$= \langle v_{n_1} \dots v_{n_k} \cos(n_1\Psi_{n_1} + \dots + n_k\Psi_{n_k}) \rangle$$

$$\langle \cos(4\varphi_1 - 2\varphi_2 - 2\varphi_3) \rangle = \langle v_4 v_2^2 \cos(4\Psi_4 - 4\Psi_2) \rangle, \langle \cos(6\varphi_1 - 3\varphi_2 - 3\varphi_3) \rangle = \langle v_6 v_3^2 \langle \cos(6\Psi_6 - 6\Psi_3) \rangle$$